

# Hyper-Kamiokande detector design and physics potential

Masashi Yokoyama

(Department of Physics, Univ. of Tokyo)

for Hyper-Kamiokande Working Group

**Fundamental Physics at the Intensity Frontier**  
November 30 - December 2, 2011  
Rockville, MD

Sep. 2011  
arXiv:1109.3262

Letter of Intent:

The Hyper-Kamiokande Experiment

— Detector Design and Physics Potential —

K. Abe,<sup>12,14</sup> T. Abe,<sup>10</sup> H. Aihara,<sup>10,14</sup> Y. Fukuda,<sup>5</sup> Y. Hayato,<sup>12,14</sup> K. Huang,<sup>4</sup>  
A. K. Ichikawa,<sup>4</sup> M. Ikeda,<sup>4</sup> K. Inoue,<sup>8,14</sup> H. Ishino,<sup>7</sup> Y. Itow,<sup>6</sup> T. Kajita,<sup>13,14</sup> J. Kameda,<sup>12,14</sup>  
Y. Kishimoto,<sup>12,14</sup> M. Koga,<sup>8,14</sup> Y. Koshio,<sup>12,14</sup> K. P. Lee,<sup>13</sup> A. Minamino,<sup>4</sup> M. Miura,<sup>12,14</sup>  
S. Moriyama,<sup>12,14</sup> M. Nakahata,<sup>12,14</sup> K. Nakamura,<sup>2,14</sup> T. Nakaya,<sup>4,14</sup> S. Nakayama,<sup>12,14</sup>  
K. Nishijima,<sup>9</sup> Y. Nishimura,<sup>12</sup> Y. Obayashi,<sup>12,14</sup> K. Okumura,<sup>13</sup> M. Sakuda,<sup>7</sup> H. Sekiya,<sup>12,14</sup>  
M. Shiozawa,<sup>12,14,\*</sup> A. T. Suzuki,<sup>3</sup> Y. Suzuki,<sup>12,14</sup> A. Takeda,<sup>12,14</sup> Y. Takeuchi,<sup>3,14</sup>  
H. K. M. Tanaka,<sup>11</sup> S. Tasaka,<sup>1</sup> T. Tomura,<sup>12</sup> M. R. Vagins,<sup>14</sup> J. Wang,<sup>10</sup> and M. Yokoyama<sup>10,14</sup>

(Hyper-Kamiokande working group)

<sup>1</sup>*Gifu University, Department of Physics, Gifu, Gifu 501-1193, Japan*

<sup>2</sup>*High Energy Accelerator Research Organization (KEK), Tsukuba, Ibaraki, Japan*

<sup>3</sup>*Kobe University, Department of Physics, Kobe, Hyogo 657-8501, Japan*

<sup>4</sup>*Kyoto University, Department of Physics, Kyoto, Kyoto 606-8502, Japan*

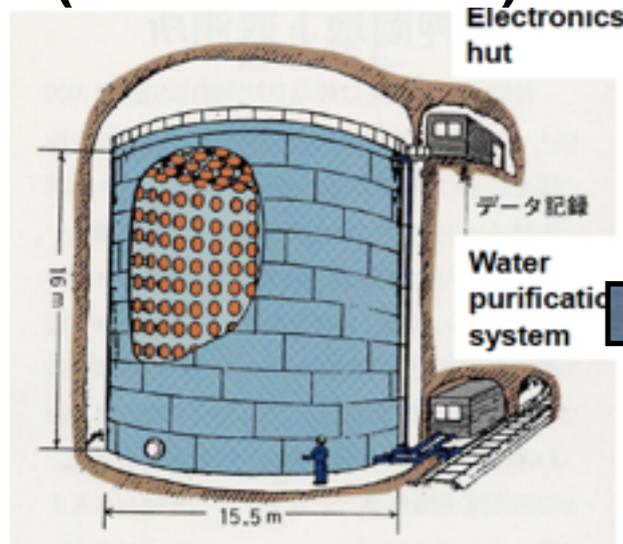
<sup>5</sup>*Miyagi University of Education, Department of Physics, Sendai, Miyagi 980-0845, Japan*

<sup>6</sup>*Nagoya University, Solar Terrestrial Environment Laboratory, Nagoya, Aichi 464-8602, Japan*

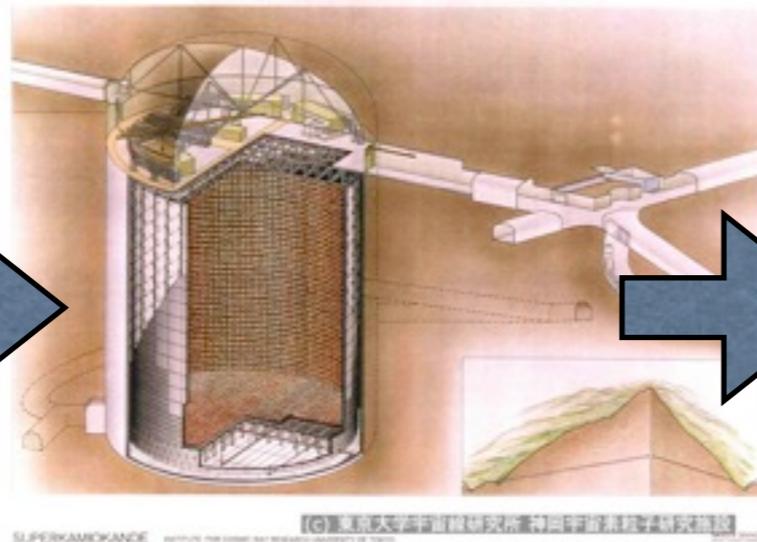
<sup>7</sup>*Okayama University, Department of Physics, Okayama, Okayama 700-8530, Japan*

# Three generations of Water Cherenkov Detectors at Kamioka

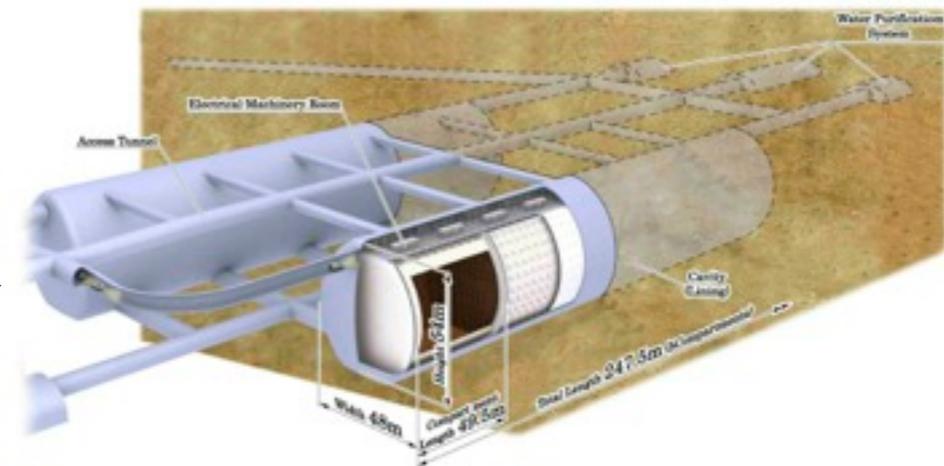
Kamiokande  
(1983-1996)



Super-Kamiokande  
(1996-)



Hyper-Kamiokande  
(201x?-)



3kton

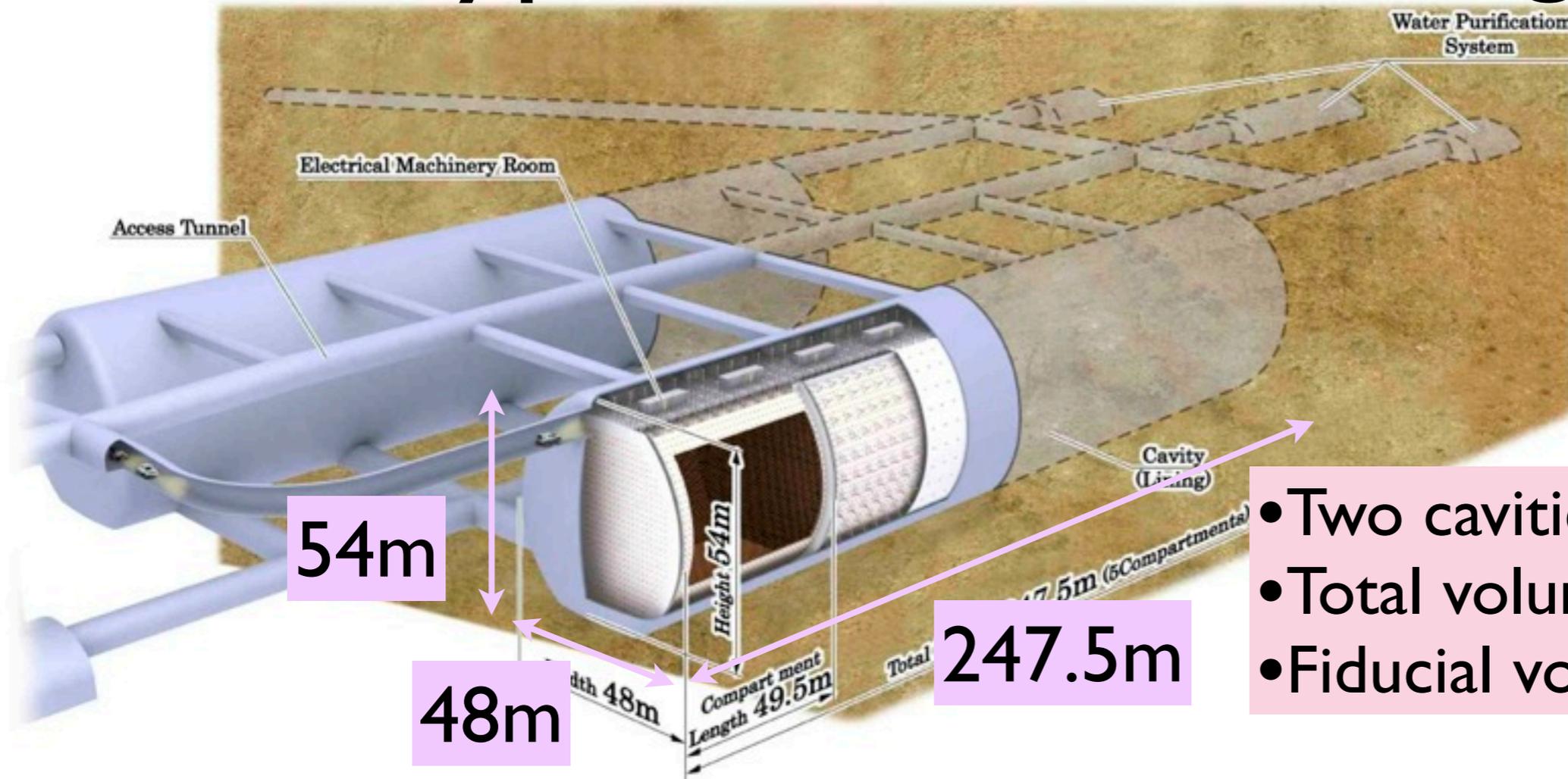
50kton

1 Mton = 1000kton

x17

x20  
(x25 fid.)

# Hyper-K: baseline design



- Two cavities
- Total volume: 0.99 Mton
- Fiducial volume: 0.56Mton

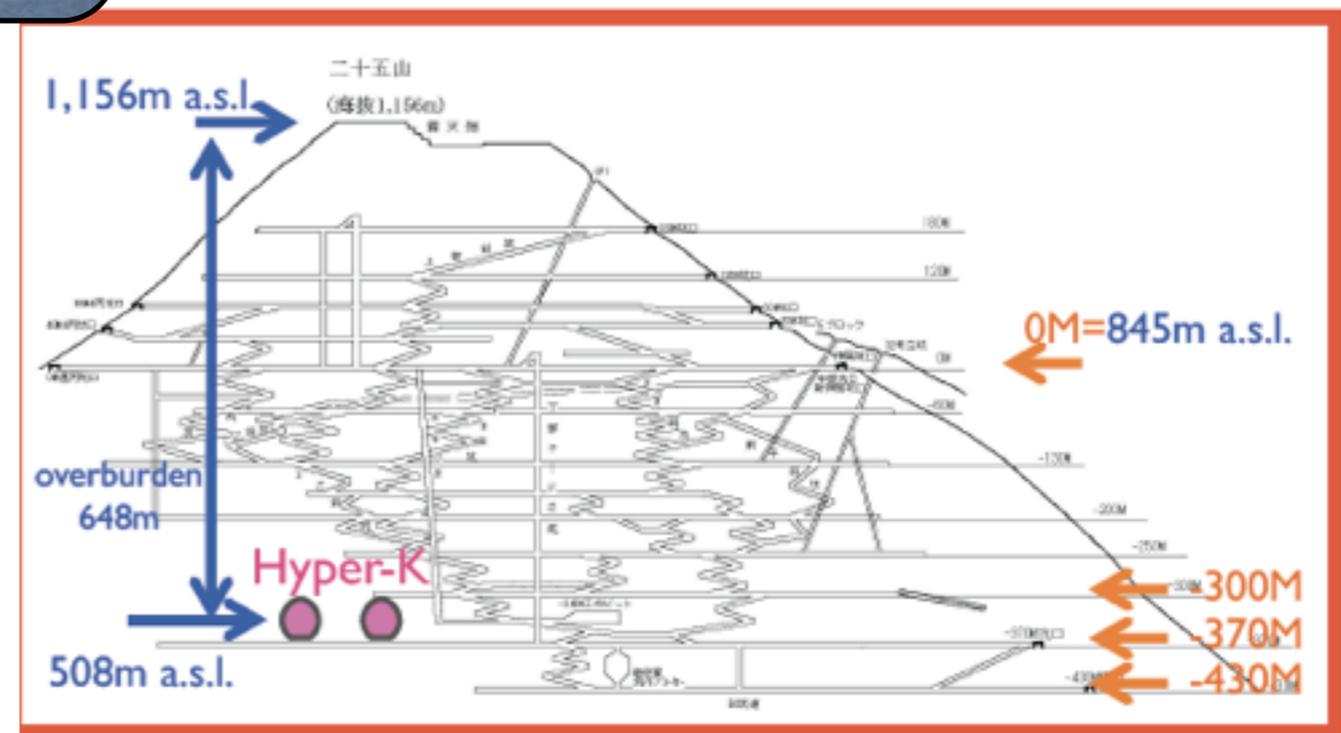
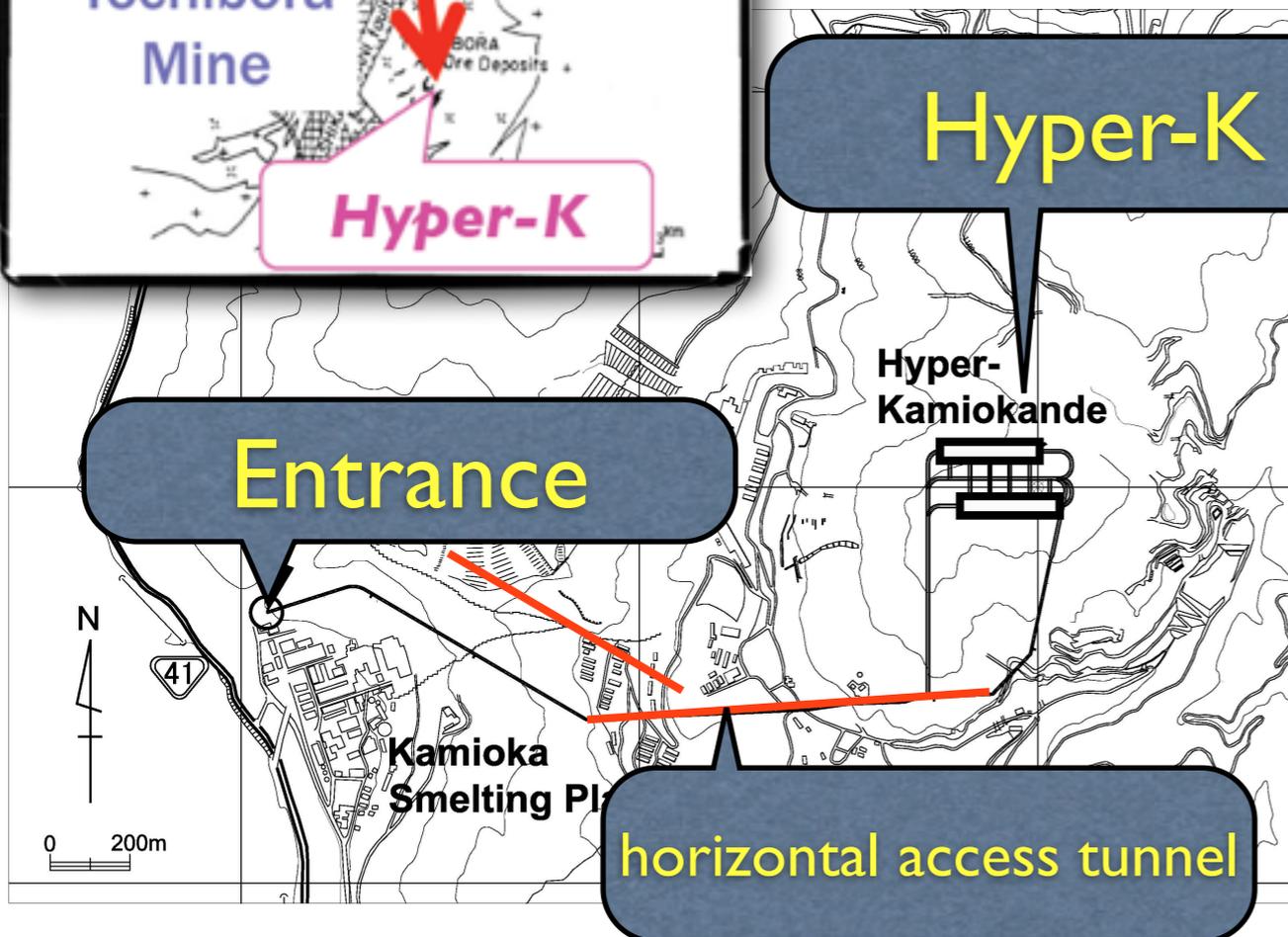
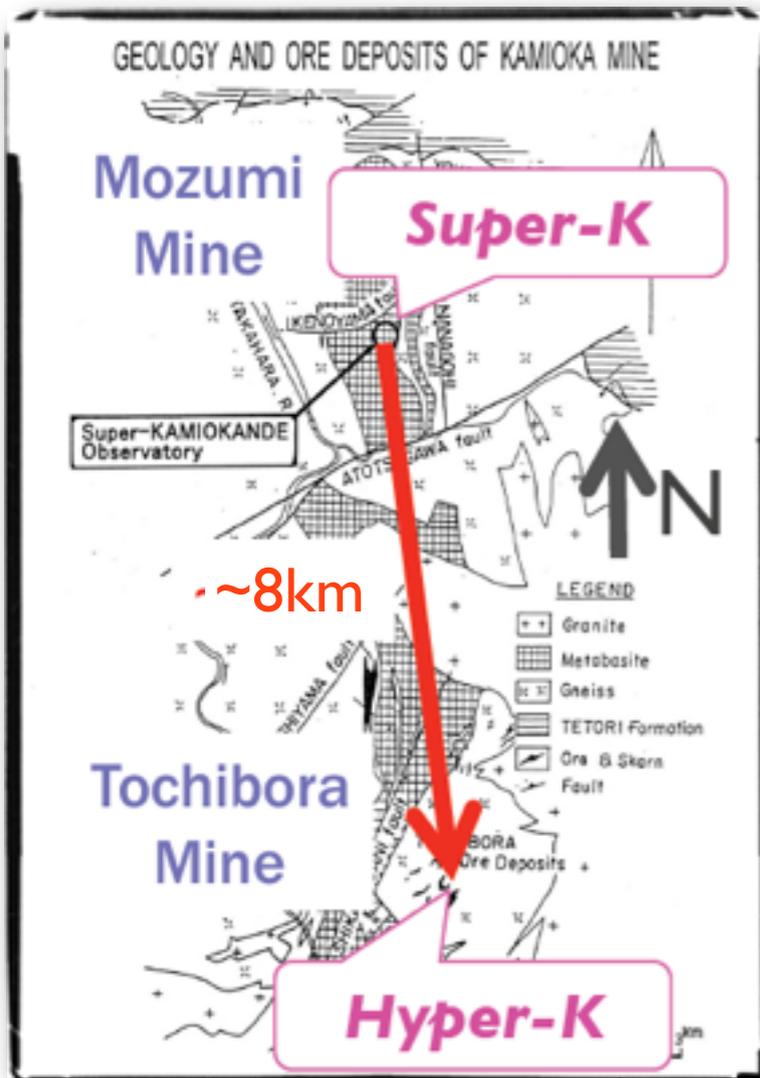
**x25 Super-K**

- Each cavity optically separated into five inner detector regions
  - Each has ~x2 volume compared to Super-K
  - Segmentation walls also instrumented
- Performance the same as Super-K

- Photo-coverage for ID: 20% (same as SK-II)
  - 99,000 20-inch PMTs
- Outer detector (2m) with 25,000 8-inch PMTs

# Candidate site

- 8km south of Super-K
- Same off-axis and baseline as T2K
- 648m of rock (1750m.w.e.) overburden
- 13,000m<sup>3</sup>/day natural water (1Mt/80days)

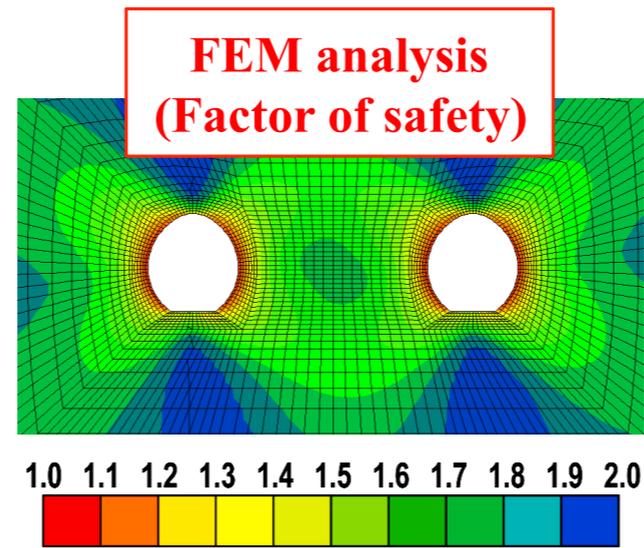


# Cavity design based on measurements

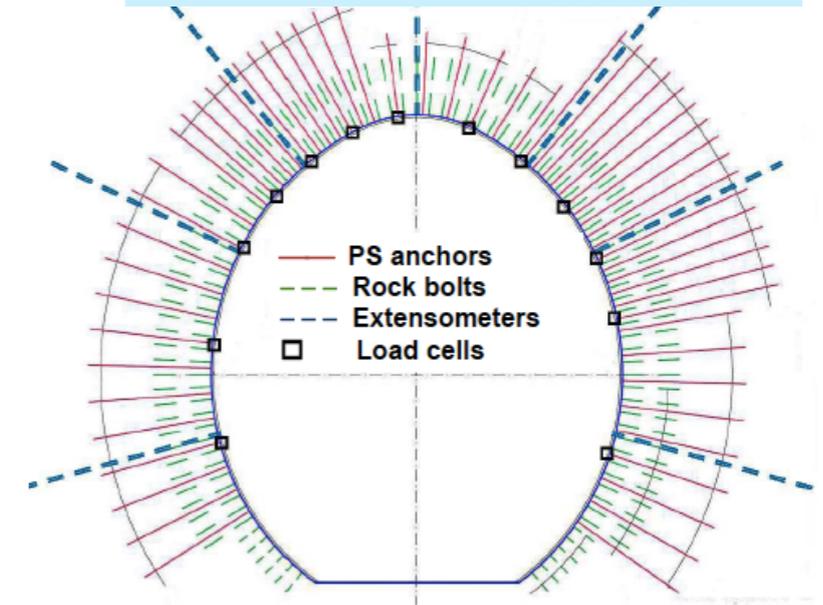
Geological survey



Twin cavern analysis

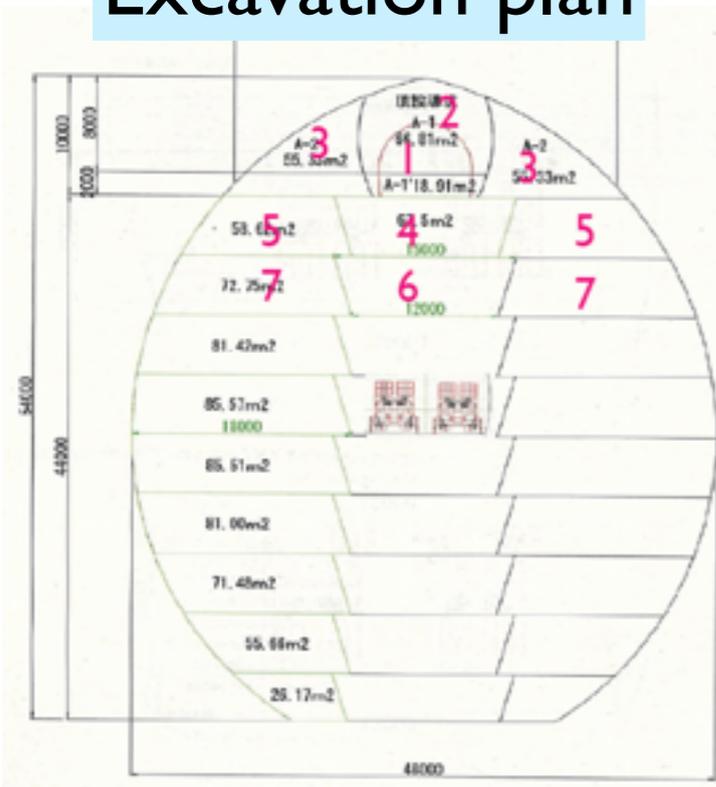


anchors, rock bolts design

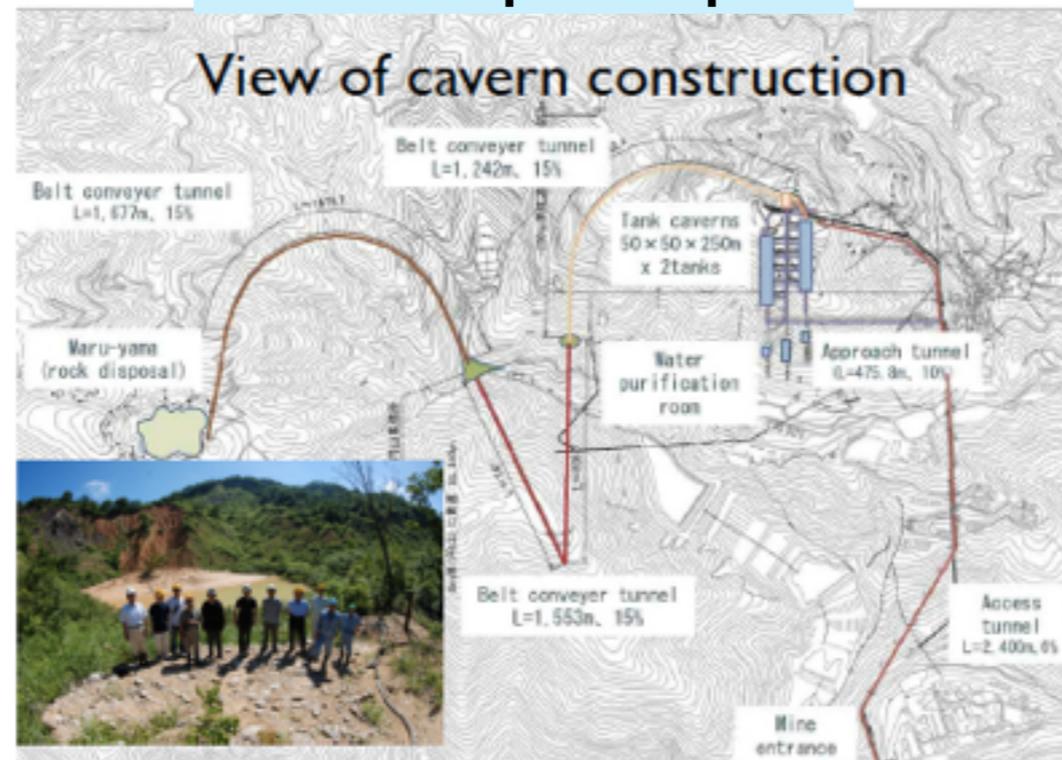


# Construction procedure

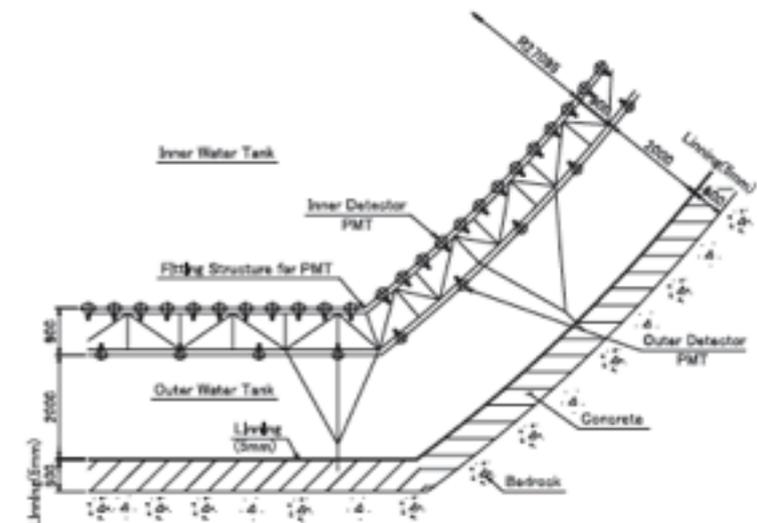
Excavation plan



Rock disposal plan



Concrete layer, polyethylene lining, PMT support



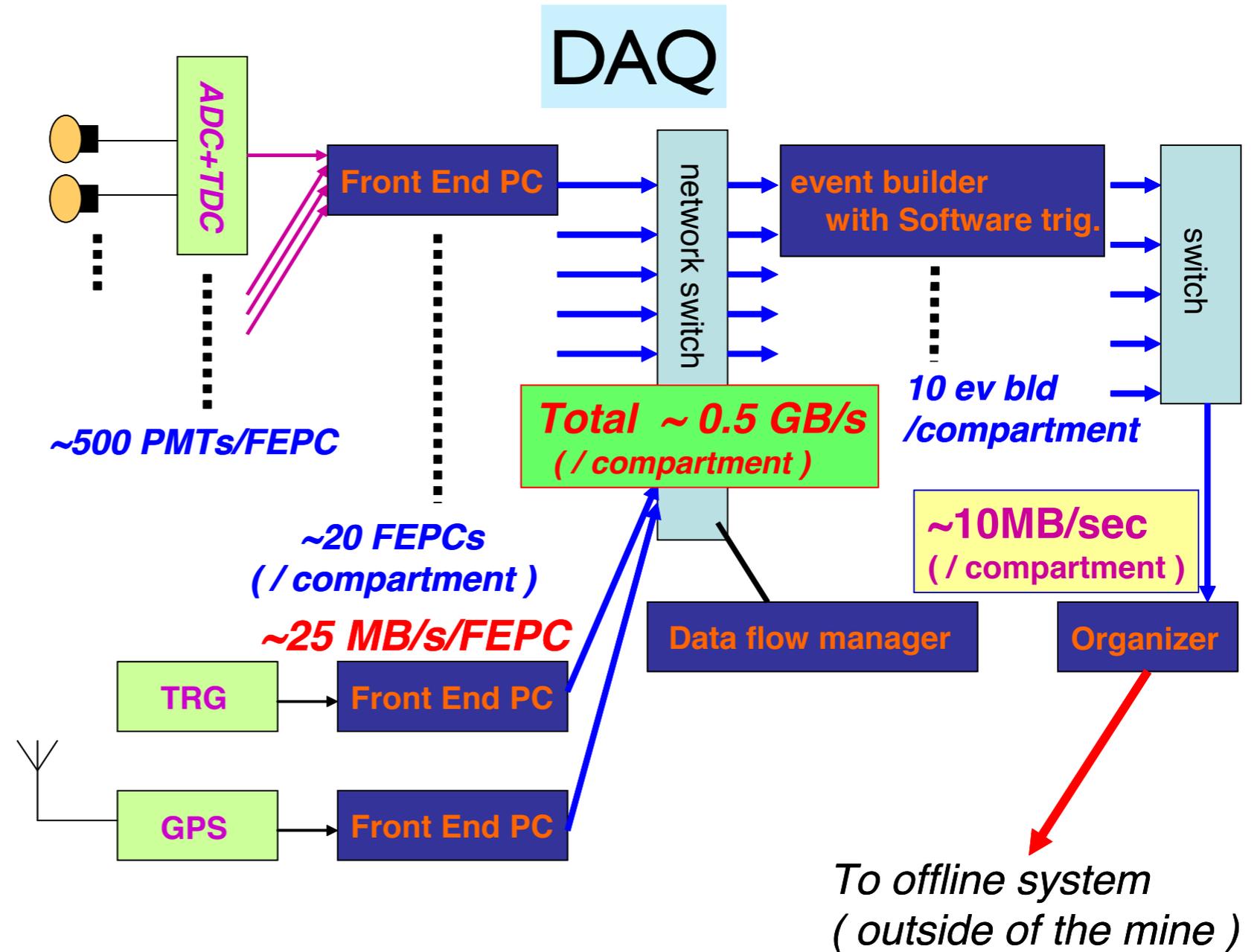
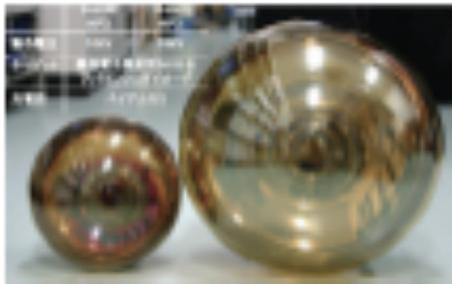
# Baseline design based on experience and R&D

## Photon sensor



8" and 13" HPDs available in 2012

• Hamamatsu will release in 2012



Detail scheduling / cost estimate on all elements underway  
More R&D effort being planned/starting

# Neutrino physics with Hyper-K

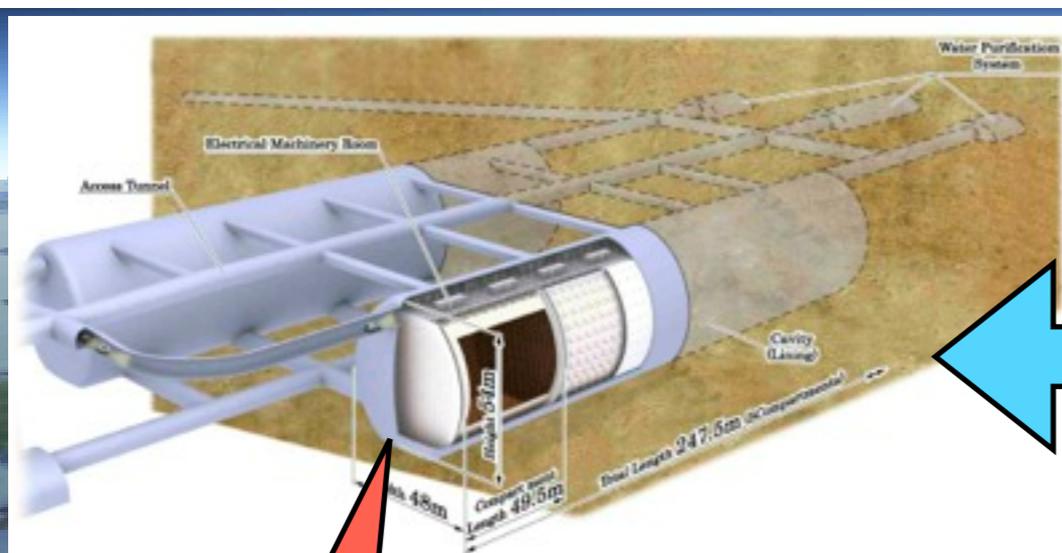
With “large”  $\theta_{13}$  (indicated by T2K, MINOS, DC, solar, ..)

Access to

- Leptonic  $CP$  violation
- $\nu$  mass hierarchy
- Octant of  $\theta_{23}$

Synergy of  
*accelerator and atmospheric*  
 $\nu$  measurements

Explore full picture of neutrino oscillation



Accelerator  $\nu$ :  
J-PARC to HK

x25 Larger Target

Quest for CP Violation  
in lepton sector



$\sim 0.6\text{GeV } \nu_{\mu}$   
295km

Higher Intensity  
1.66MW (KEK roadmap)



JPARC



© 2010 ZENRIN  
Data © 2010 MIRC/JHA  
© 2010 Cnes/Spot Image  
© 2010 Mapabc.com

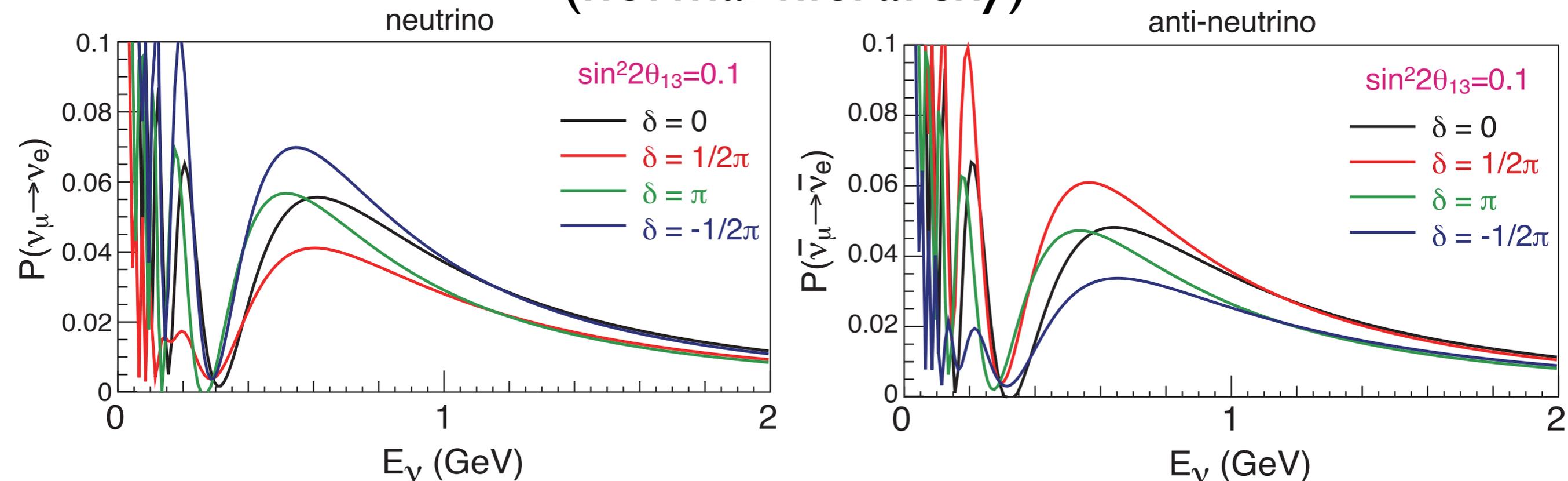
36°24'46.66" N 139°18'01.27" E 標高 214 メートル

高度 188.55 キロメートル

©2009 Google

# Measuring $CP$ asymmetry

$P(\nu_\mu \rightarrow \nu_e)$  appearance probability  
(normal hierarchy)



- Comparison of  $P(\nu_\mu \rightarrow \nu_e)$  and  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$
- In principle, sensitive to exotic (non-MNS) CPV

# Full MC study

- Full simulation of  $\nu$  beam, interaction, detector response and reconstruction
  - Number of PMT: half of SK-4 (20% coverage)
  - Event selection almost the same as T2K
    - (for  $E_{\nu}^{\text{rec}} < 1.25 \text{ GeV}$ )
  - Loose cut on  $E_{\nu}$  to utilize spectrum information
- Assumed total five year running
  - 1.5 year in  $\nu$  / 3.5 year in  $\bar{\nu}$
  - 1.66MW beam power for J-PARC assumed

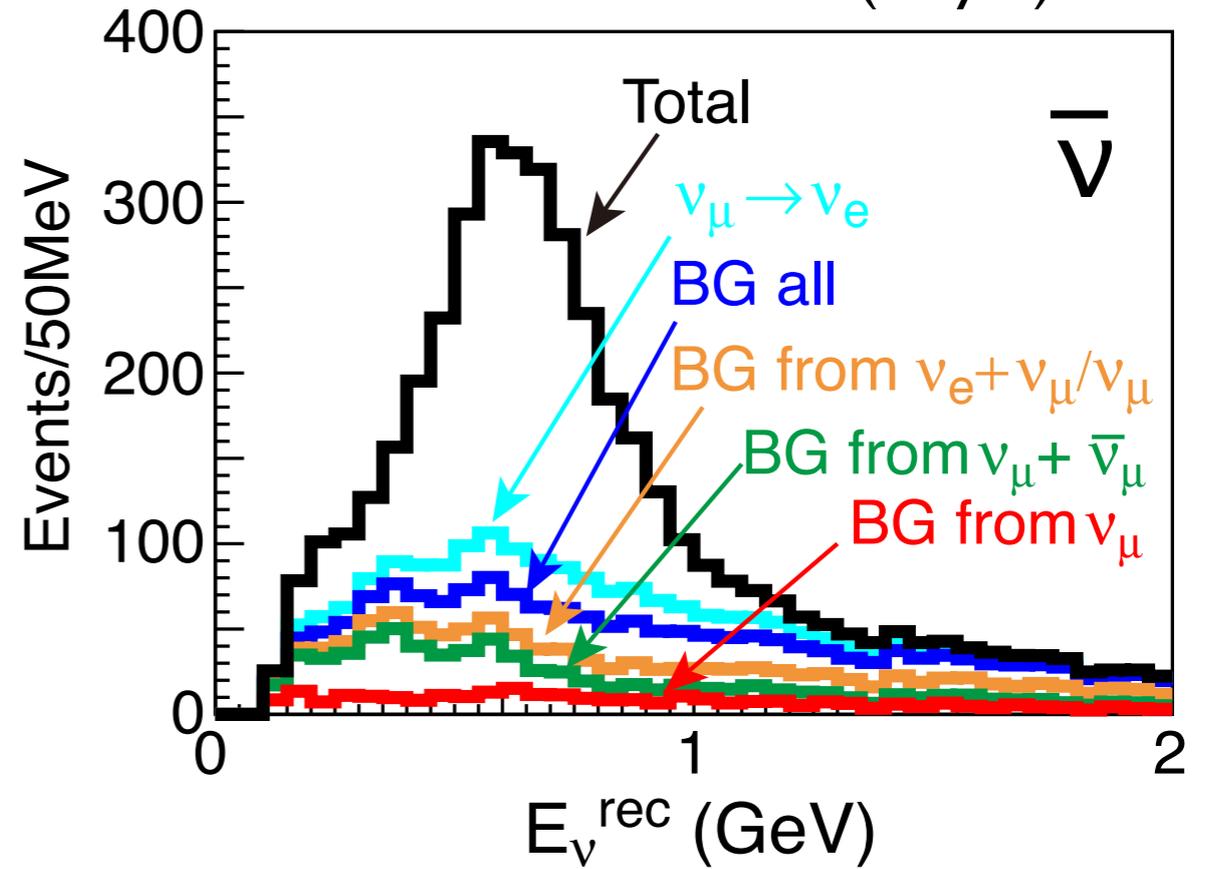
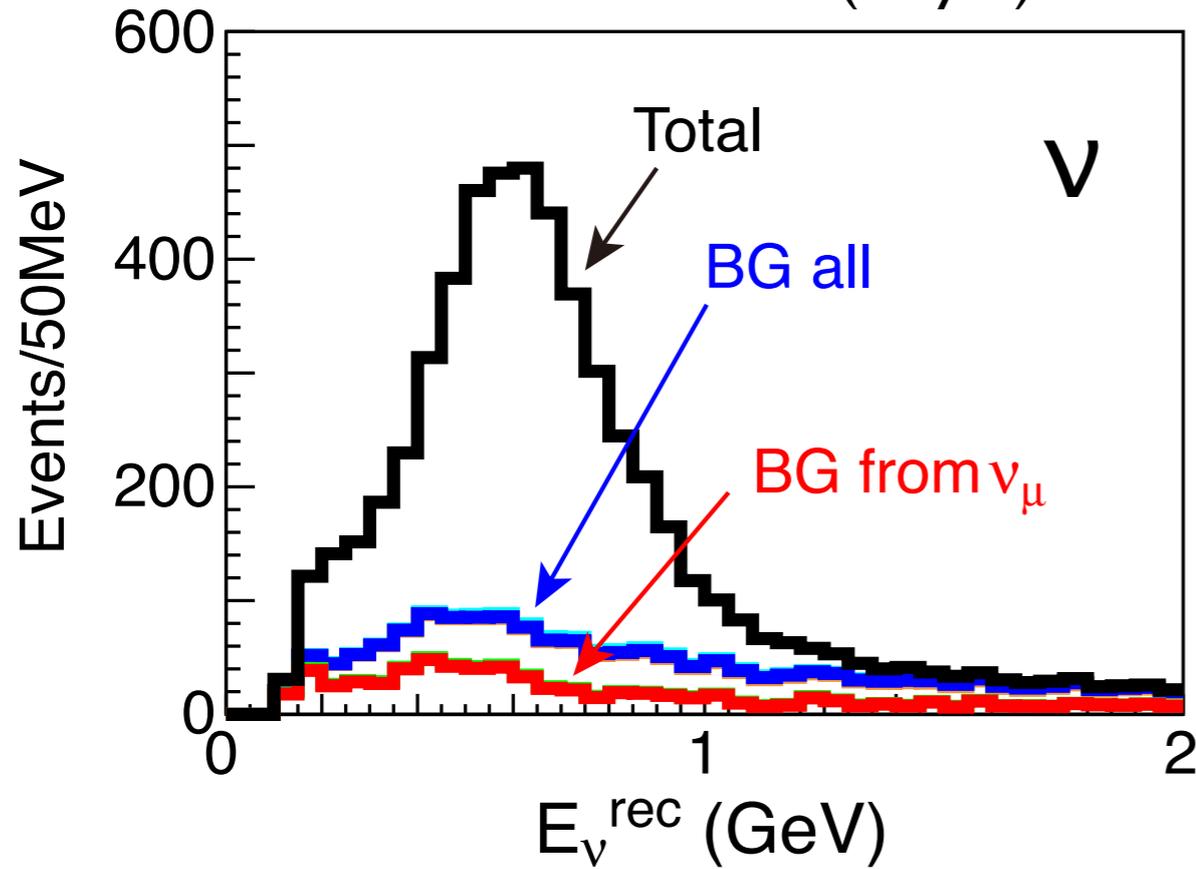
Signal efficiency	64%
$\nu_{\mu}$ CC BG rejection	>99.9%
NC $\pi^0$ BG rejection	95%

# $\nu_e$ candidates after selection

$\sin^2 2\theta_{13} = 0.1, \delta = 0$

$\nu$  mode (1.5yrs)

$\bar{\nu}$  mode (3.5yrs)



	Signal ( $\nu_\mu \rightarrow \nu_e$ CC)	$\nu_\mu / \bar{\nu}_\mu$ CC	$\nu_e / \bar{\nu}_e$ CC	NC
$\nu$ (1.5yrs)	<b>3,991</b>	39	974	718
$\bar{\nu}$ (3.5yrs)	<b>2,589</b>	25	972	750

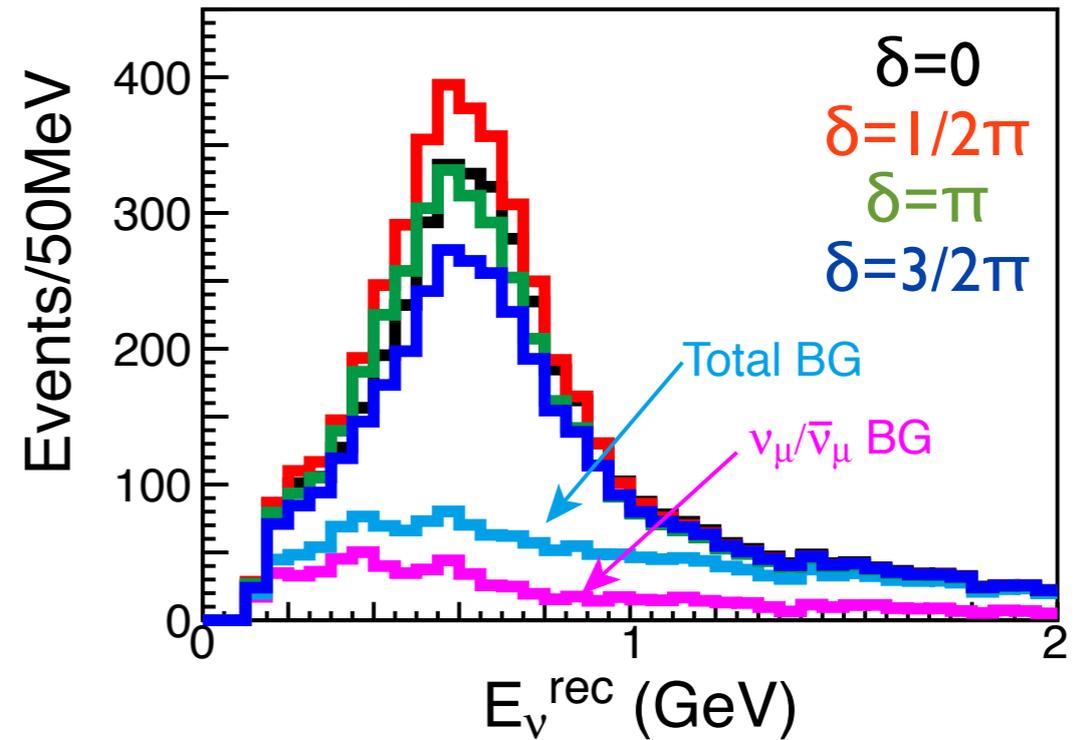
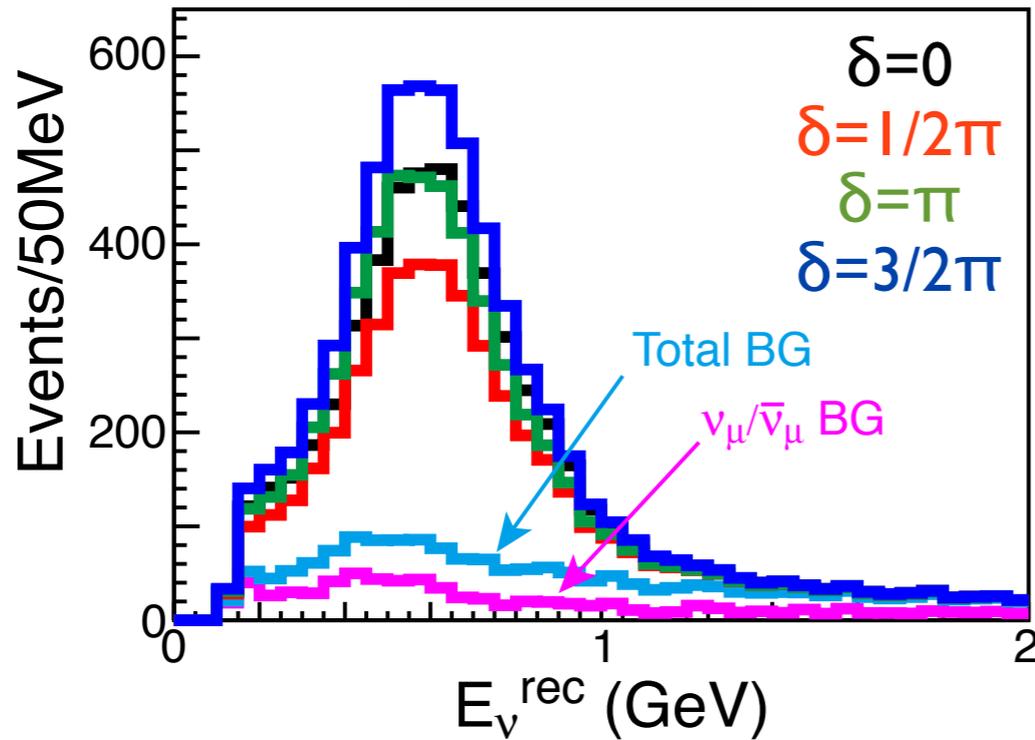
( $E_\nu^{\text{rec}} < 2\text{GeV}$ )

$$\sin^2 2\theta_{13} = 0.1$$

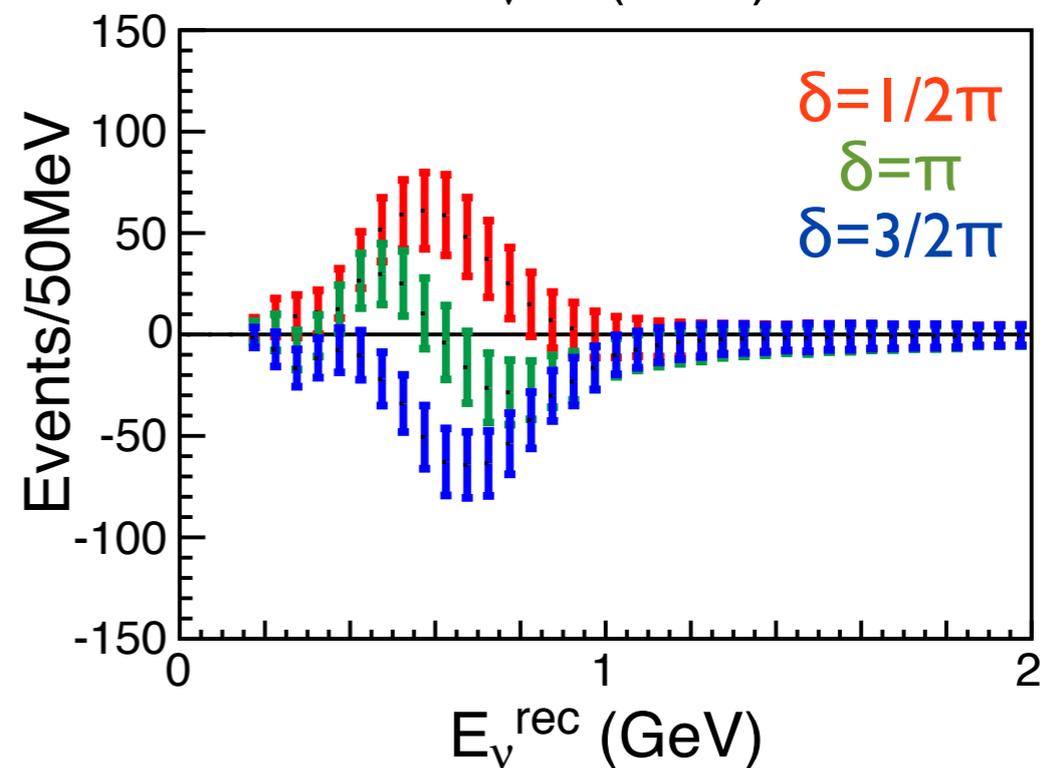
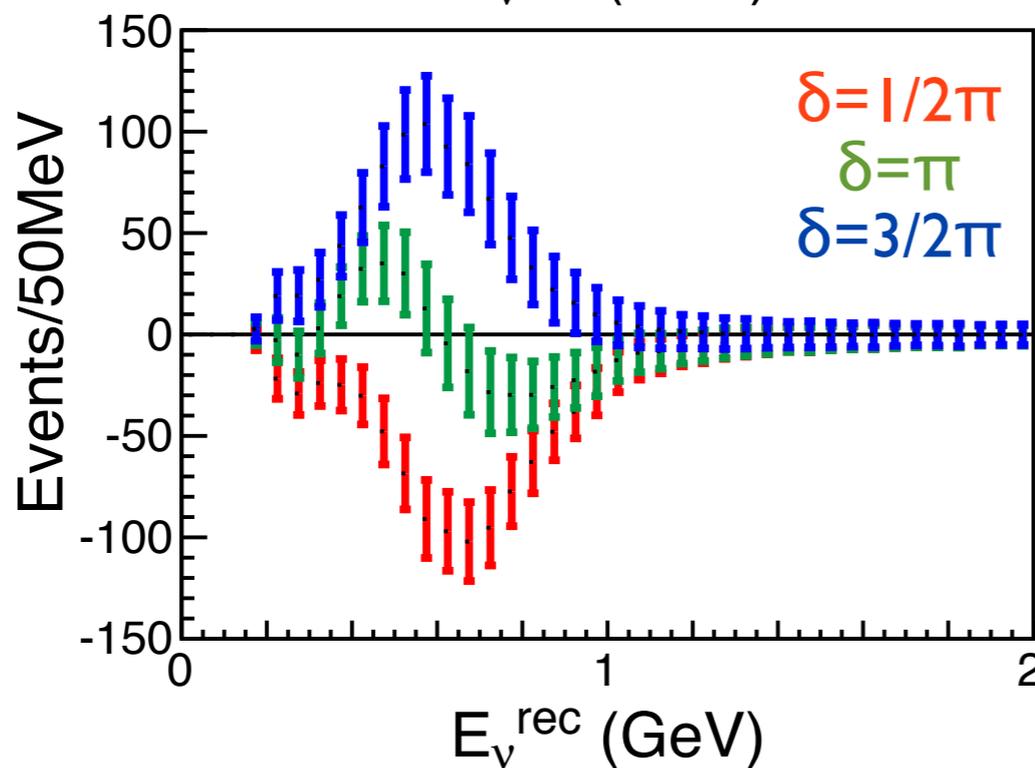
$\nu$  1.5yrs

$\bar{\nu}$  3.5yrs

ve candidates



Difference from  $\delta=0$

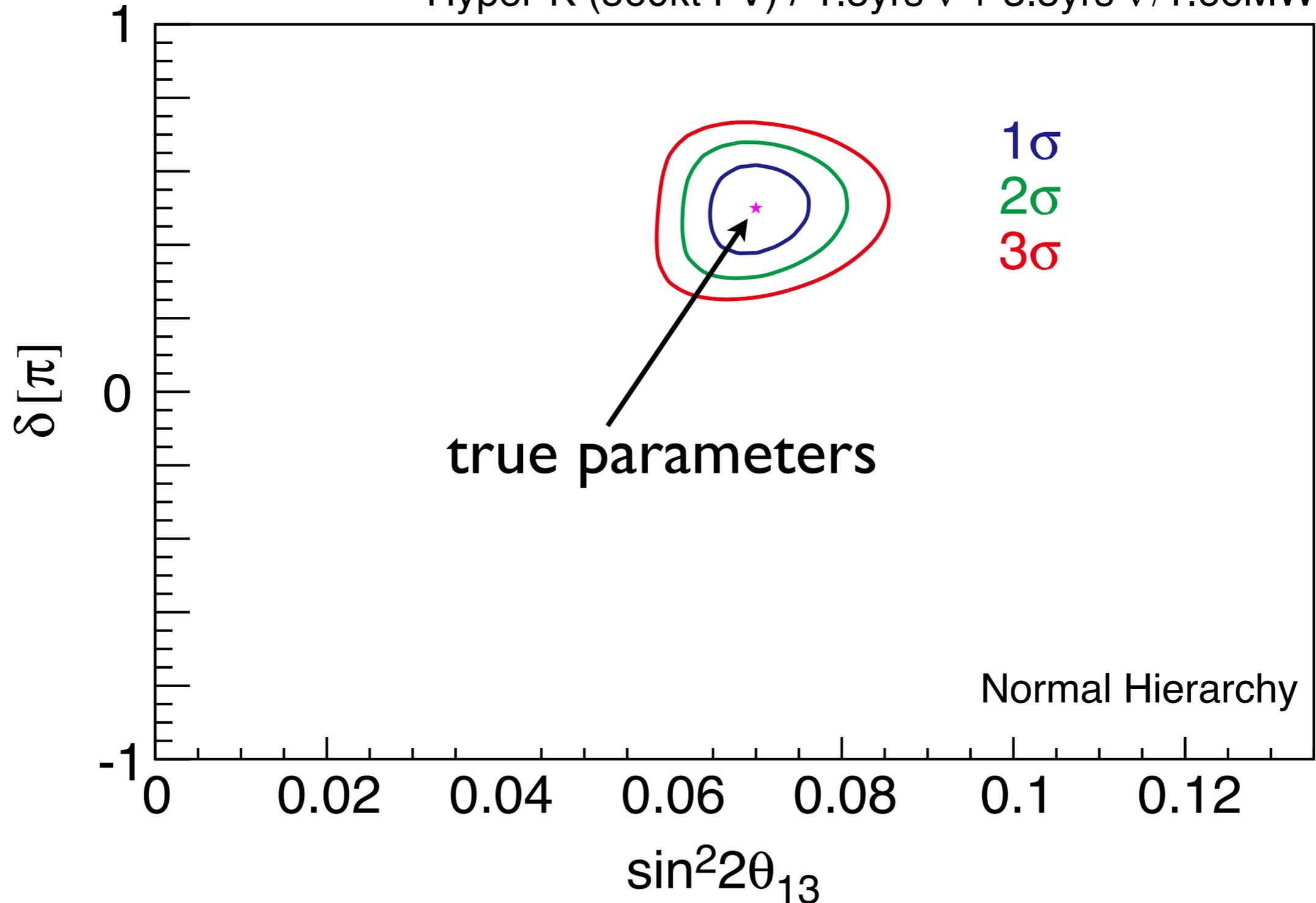


Sensitive to all values of  $\delta$  (including  $\delta=0, \pi$ )

# Expected contour

5% systematics on signal,  $\nu_\mu$  BG,  $\nu_e$  BG,  $\nu/\bar{\nu}$

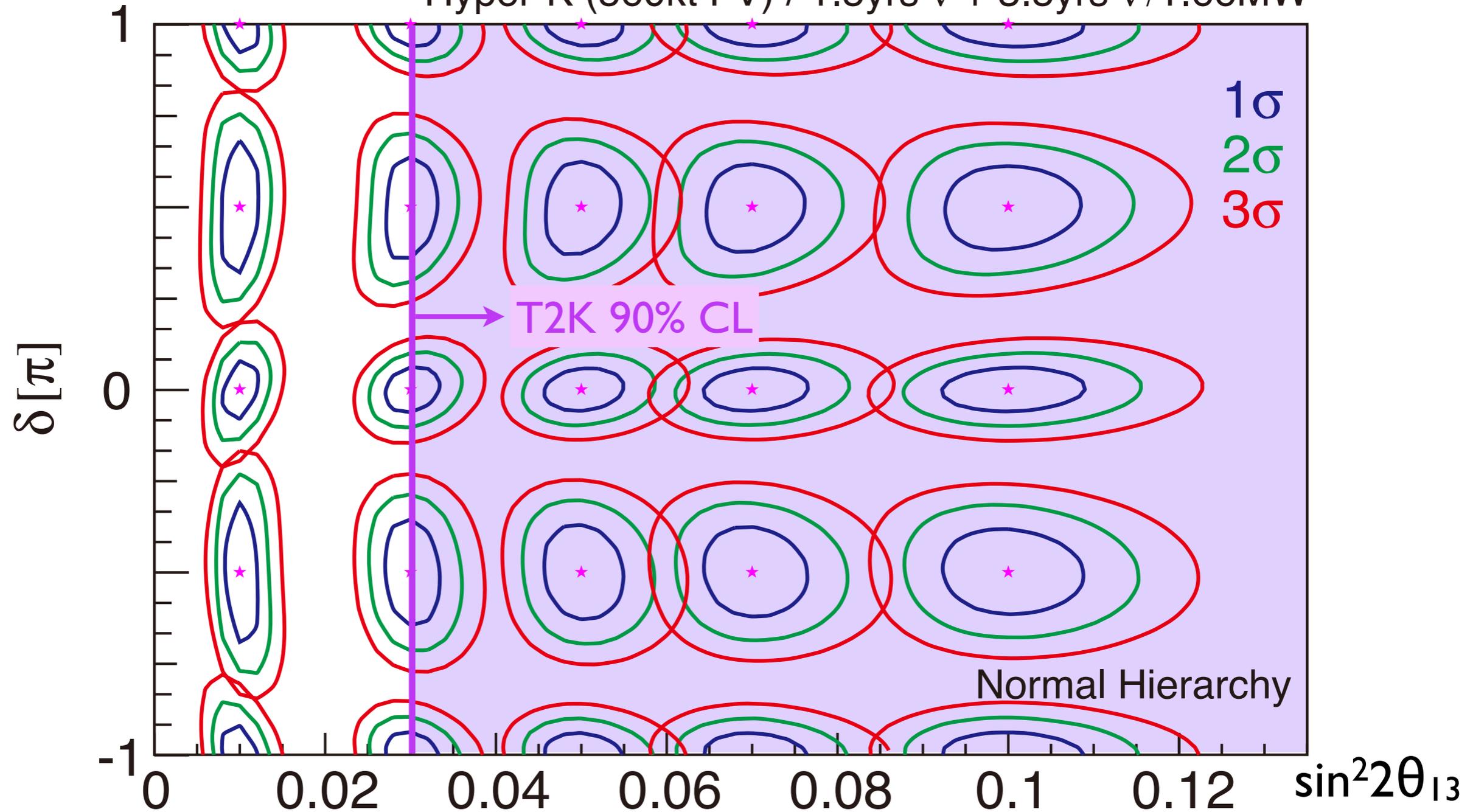
Hyper-K (560kt FV) / 1.5yrs  $\nu$  + 3.5yrs  $\bar{\nu}$  / 1.66MW



# Expected contours

5% systematics on signal,  $\nu_\mu$  BG,  $\nu_e$  BG,  $\nu/\bar{\nu}$

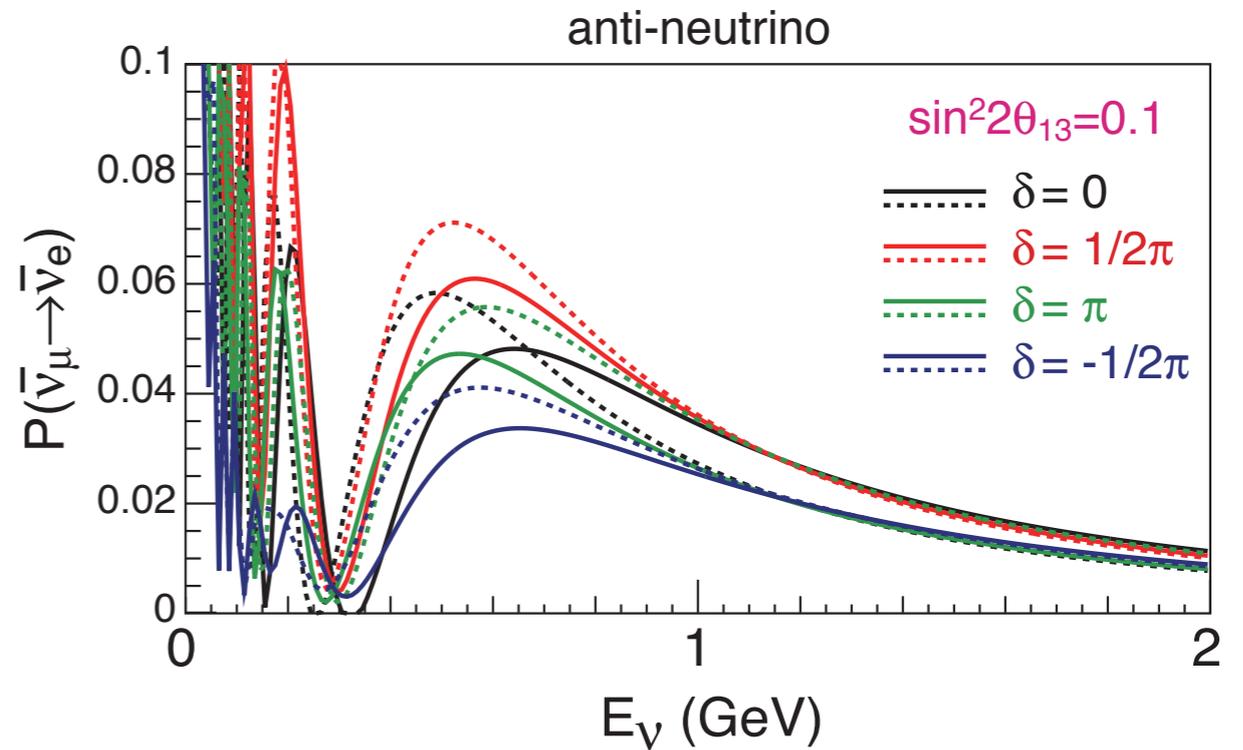
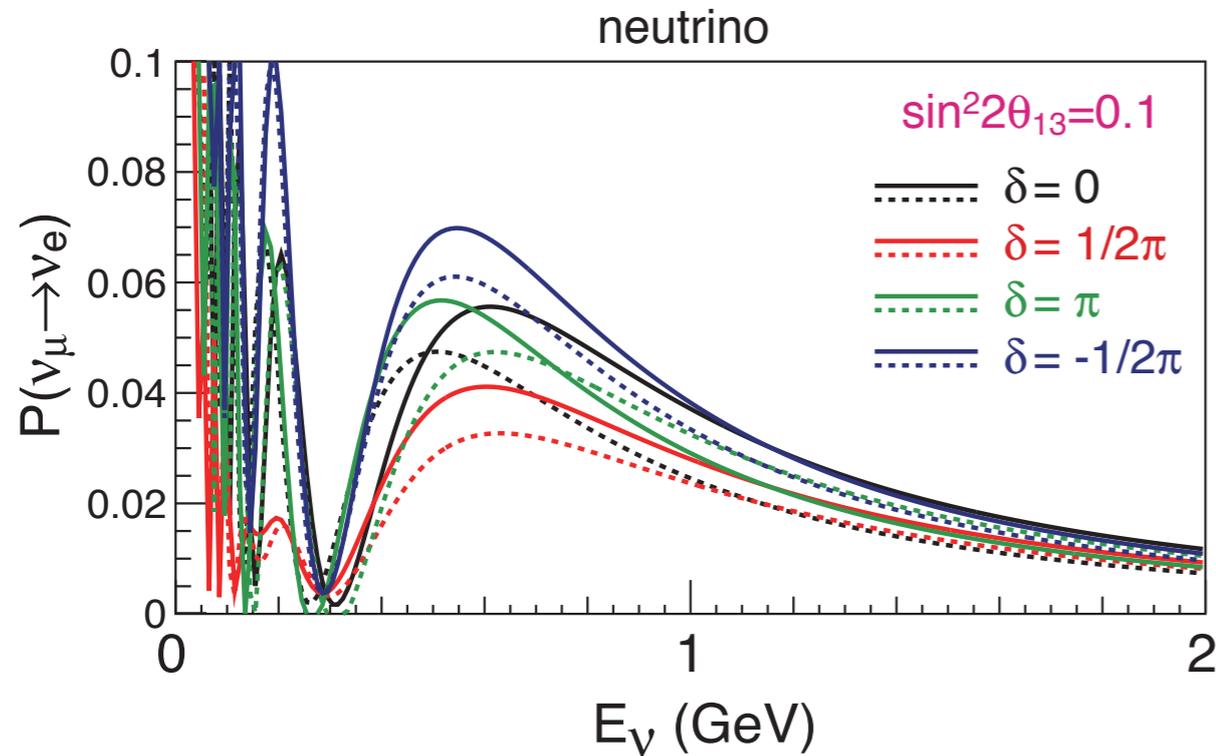
Hyper-K (560kt FV) / 1.5yrs  $\nu$  + 3.5yrs  $\bar{\nu}$  / 1.66MW



Good sensitivity over allowed region of T2K

# Mass hierarchy

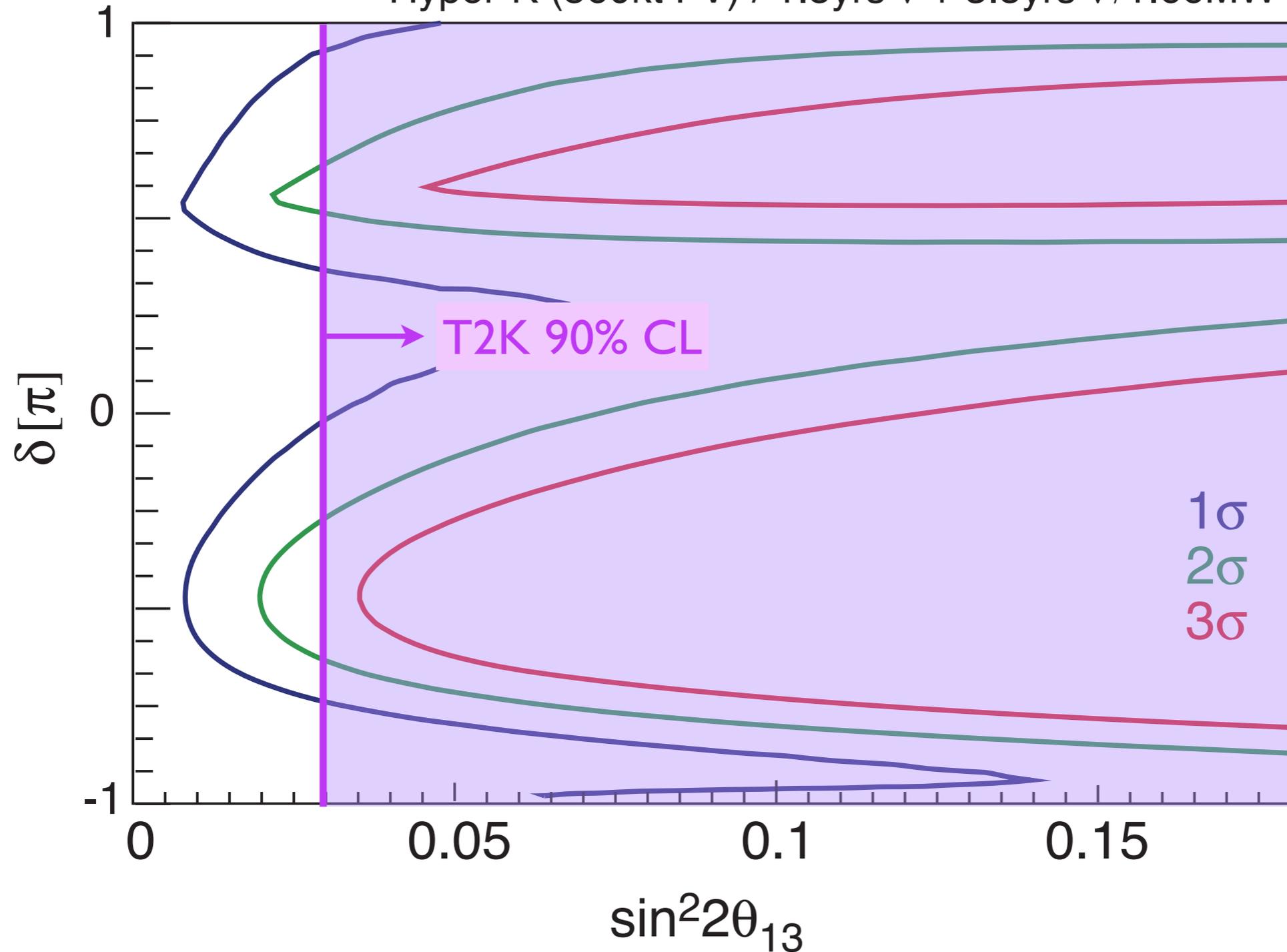
Solid: normal  
Dashed: inverted



- Fake CPV effect by matter effect
- Some sensitivity using shape information

# Mass hierarchy

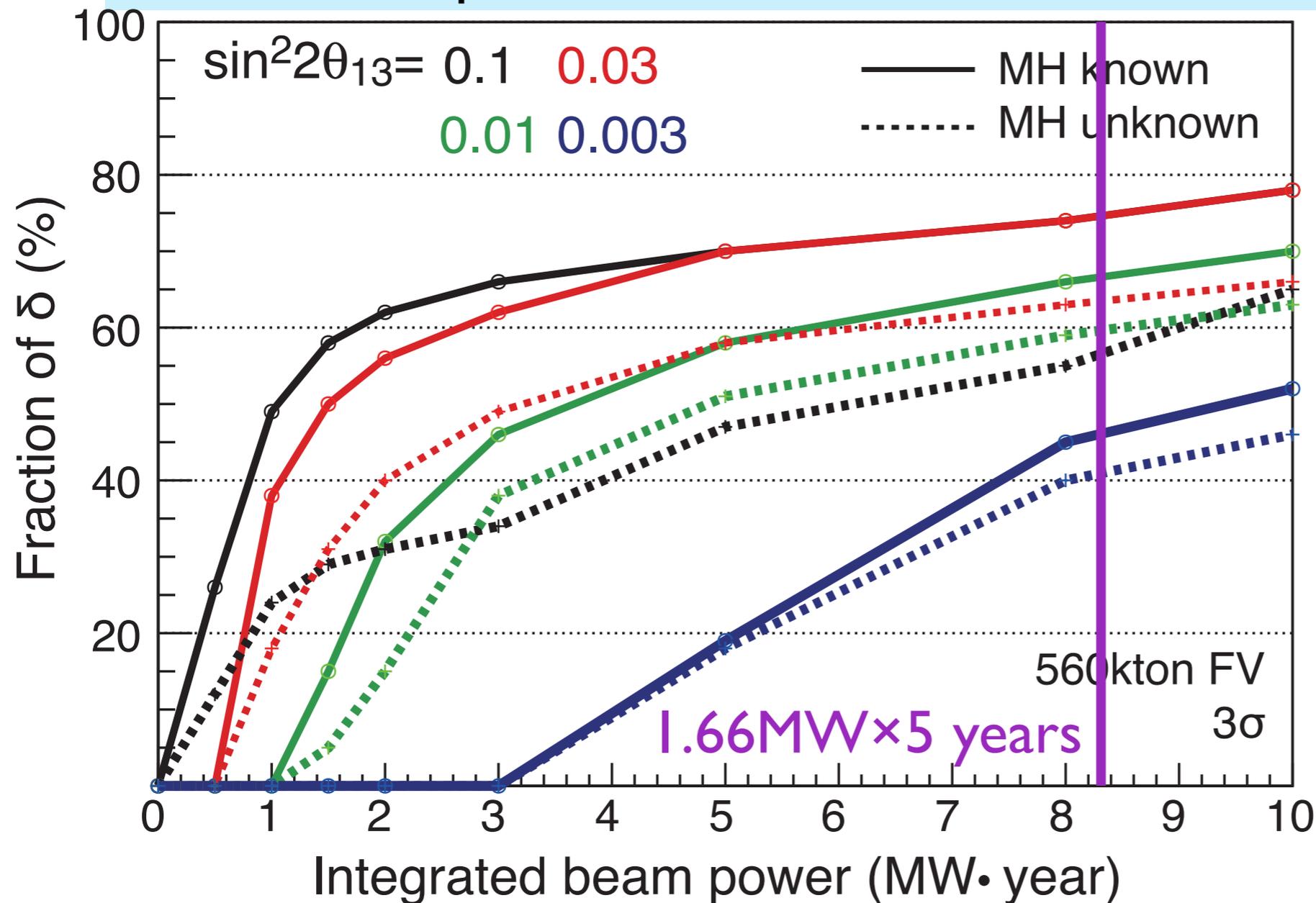
Hyper-K (560kt FV) / 1.5yrs  $\nu$  + 3.5yrs  $\bar{\nu}$  / 1.66MW



J-PARC HK experiment has sensitivity to mass hierarchy

# CPV sensitivity

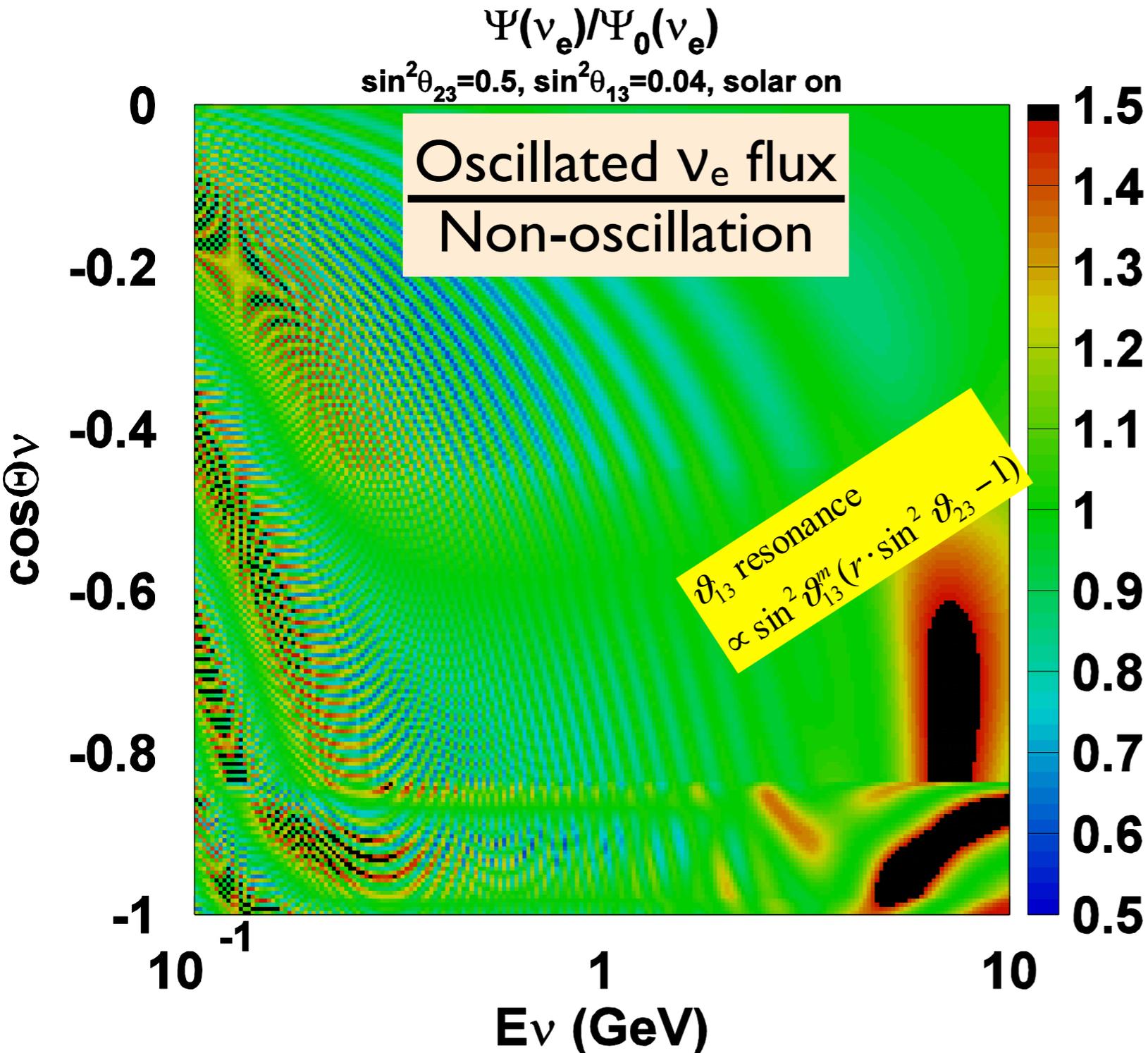
Fraction of  $\delta$  parameter for which CPV can be observed with  $>3\sigma$



$\sin^2 2\theta_{13}$	Mass hierarchy	
	known	unknown
0.1	74	55
0.03	74	63
0.01	66	59

- CPV can be observed with  $>3\sigma$  for 74% of  $\delta$  ( $\sin^2 2\theta_{13} > 0.03$ )
- Effect of unknown mass hierarchy is limited
- Input from atm  $\nu$  and other experiments also expected for MH

# Atmospheric $\nu$



$\nu_\mu \rightarrow \nu_e$  appearance  
 resonance either  $\nu$  or  $\bar{\nu}$   
 depending on  
 mass hierarchy

Sensitive to

- Mass hierarchy
- $\theta_{23}$  octant
- CP asymmetry

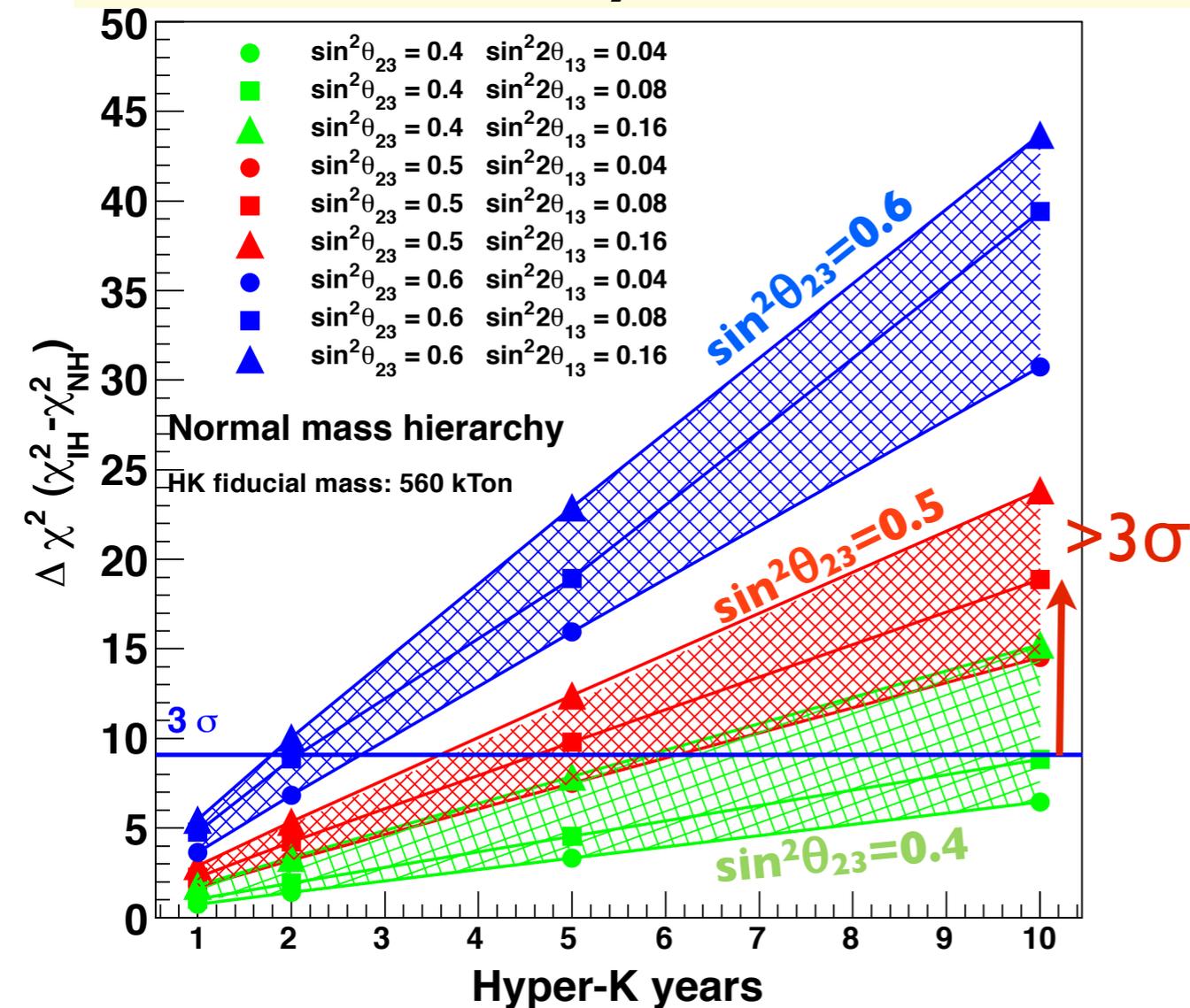
Complementary to  
 accelerator  $\nu$

Large  $\theta_{13}$  gives  
 better sensitivity

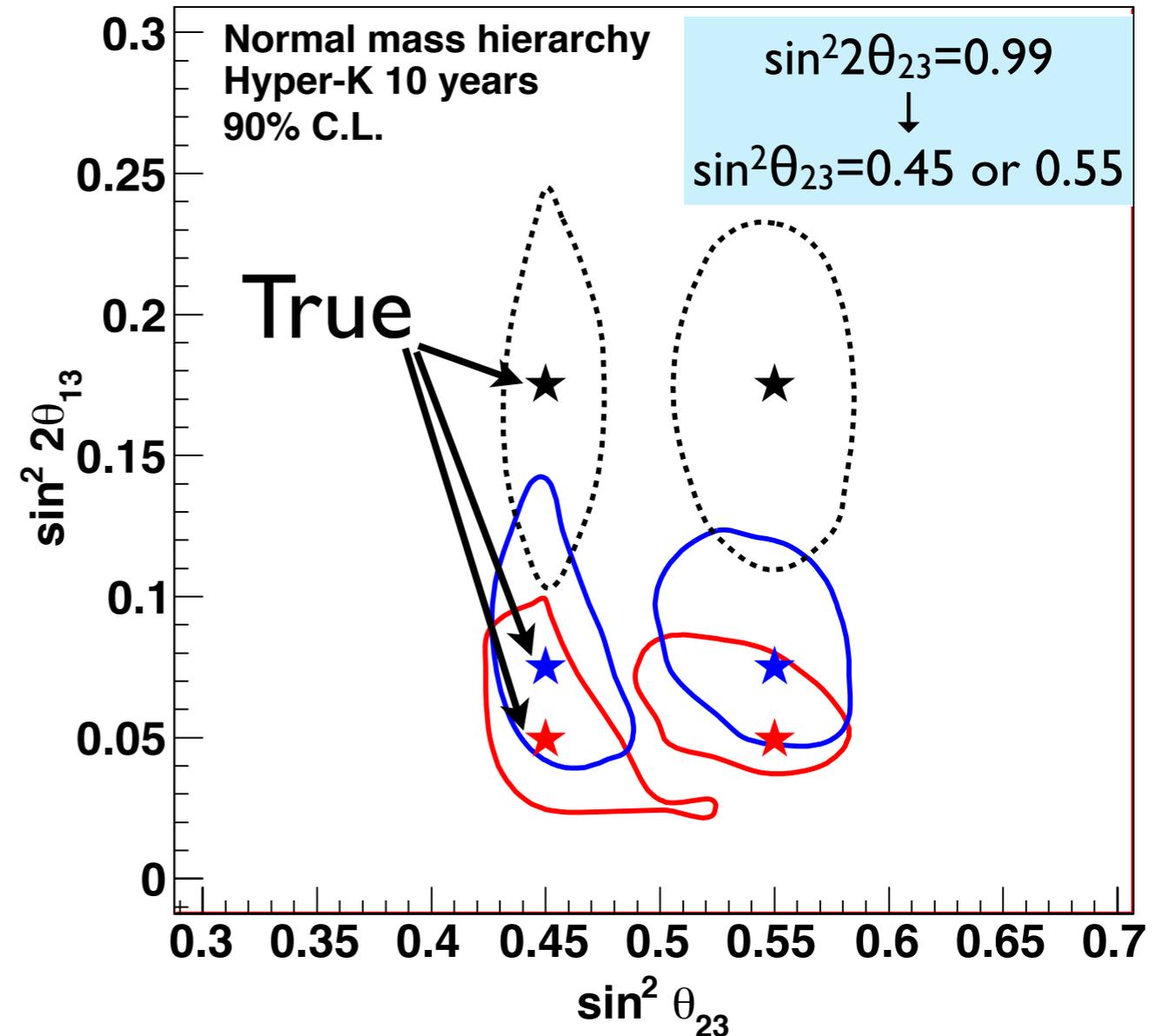
# Atmospheric $\nu$

## Mass hierarchy determination

## $\theta_{23}$ octant



$> 3\sigma$  in 5-10 years

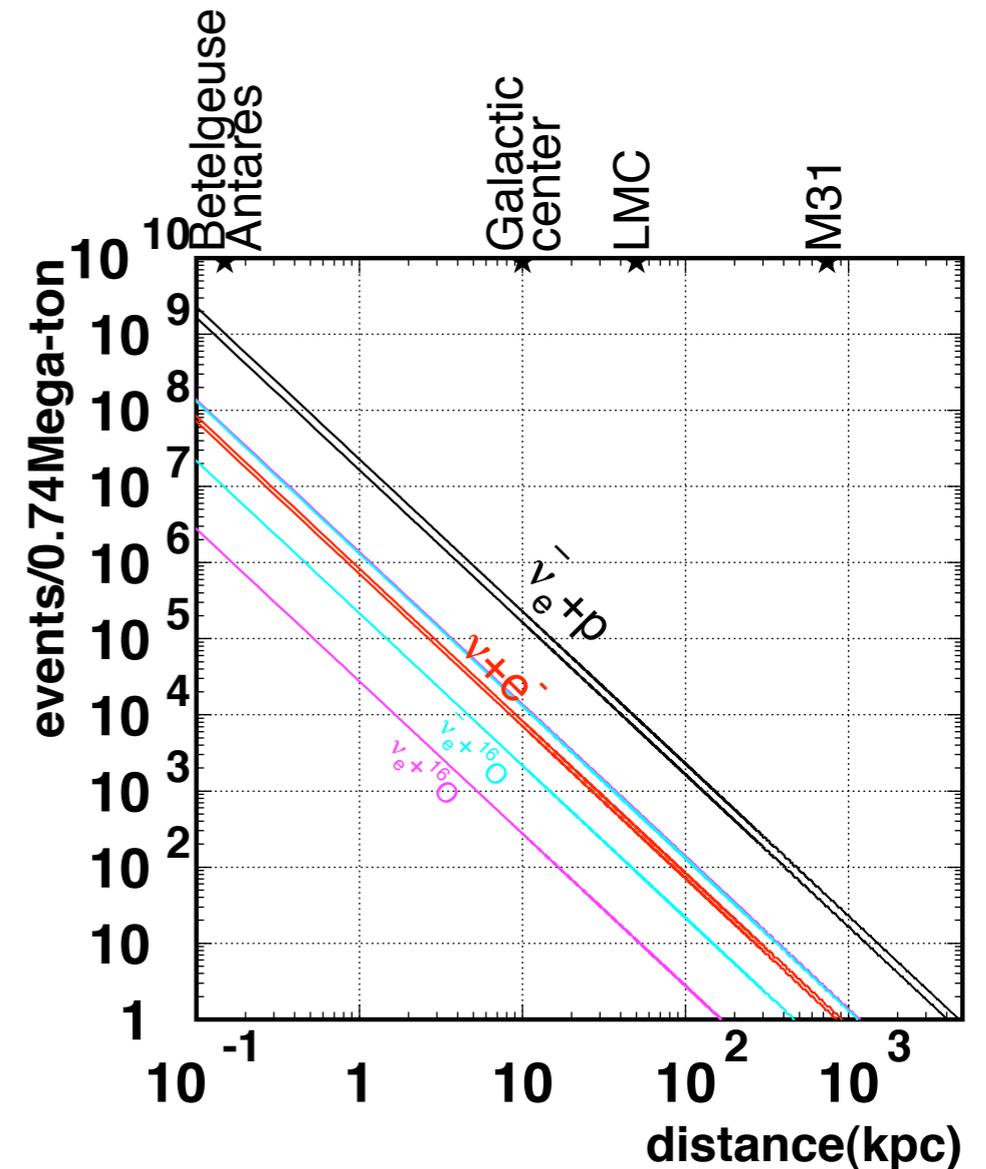


Resolved if  $\sin^2 2\theta_{23} < 0.99$

Complementary to accelerator  $\nu$

# Other science with $\nu$ 's in HK

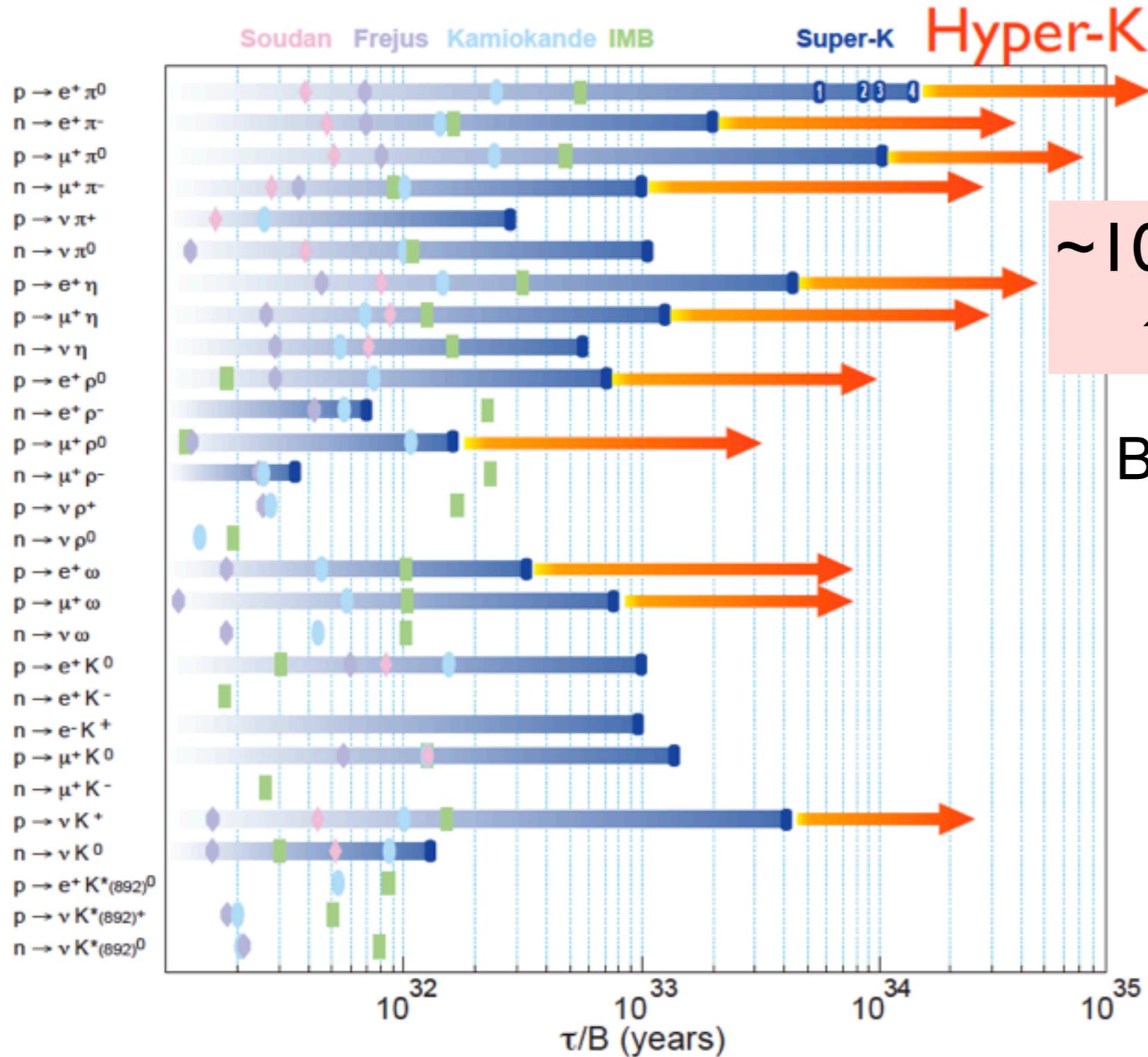
- Supernova core collapse  $\nu$ 
  - $\sim 200\text{k}$  events for 10kpc SN
- Relic supernova  $\nu$  (with Gd?)
- Solar  $\nu$  precise measurements
- WIMP, GRB, solar flare...
- Geophysics ( $\nu$  tomography of Earth)



covering a wide range of particle physics/astrophysics topics

# Nucleon decay with HK

Talk by Y. Obayashi  
in nucleon decay session



~10 times better sensitivity  
than current Super-K

Based on full simulation

- $p \rightarrow e^+ \pi^0$ :
  - $1.3 \times 10^{35}$  yrs (90%CL)
  - $5.7 \times 10^{34}$  yrs ( $3\sigma$ )
- $p \rightarrow \nu K^+$ :
  - $2.5 \times 10^{34}$  yrs (90%CL)
  - $1.0 \times 10^{34}$  yrs ( $3\sigma$ )

(10 years)

# Summary

- Baseline design and physics potential of Hyper-Kamiokande are presented. [arXiv:1109.3262]
- excellent sensitivity to leptonic CP violation.
- discrimination of mass hierarchy and  $\theta_{23}$  octant.
- order of magnitude better sensitivity for nucleon decays.
- astrophysical  $\nu$  observation.
- Next step:
  - Detailed, realistic cost and schedule.
  - Optimization of design, more R&D.

**Backup**

# $\nu$ physics case of Hyper-K

Assuming  $\sin^2\theta_{13}\sim$ a few % (suggested by T2K, MINOS, DC)

- ▶ Leptonic CP violation, Dirac phase  $\delta$
- ▶  $\nu$  mass hierarchy,  $\Delta m^2_{32}>0$  or  $\Delta m^2_{32}<0$
- ▶  $\theta_{23}$  octant,  $\theta_{23}<\pi/4$  or  $\theta_{23}>\pi/4$

Proposal to explore full picture of neutrino oscillation parameters.

Proton Decay searches w/  $\sim 10$  times better sensitivity than Super-K, astrophysical  $\nu$  etc.

# Assumed parameters

- $L = 295 \text{ km}$
- $\Delta m^2_{12} = 7.6 \times 10^{-5} \text{ eV}^2$
- $|\Delta m^2_{23}| = 2.4 \times 10^{-3} \text{ eV}^2$
- $\sin^2 \theta_{12} = 0.31$
- $\sin^2 \theta_{23} = 0.50$  ( $\sin^2 2\theta_{23} = 1.0$ )
- $\rho = 2.6 \text{ [g/cm}^3\text{]}$

# $\chi^2$ definition

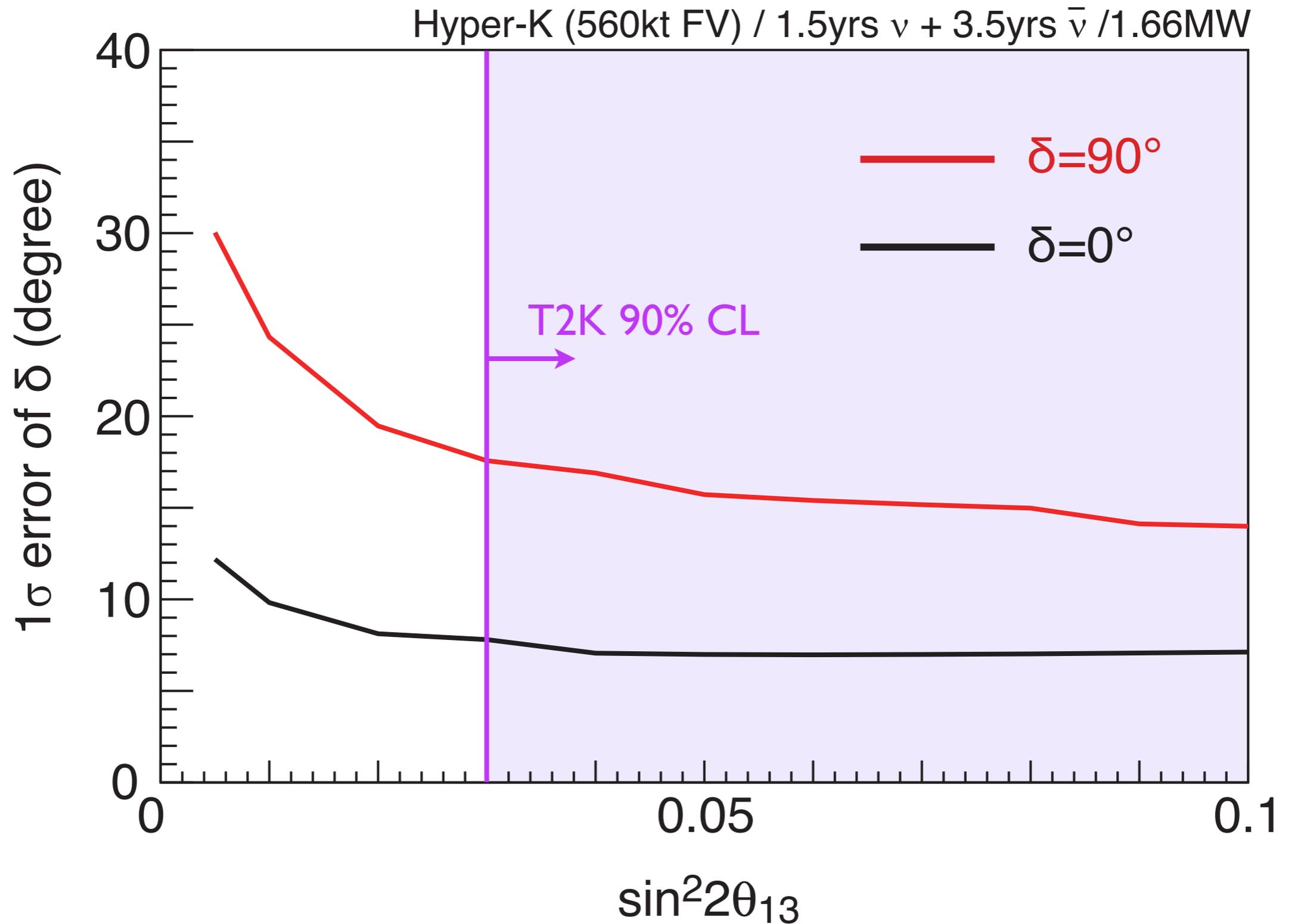
Fake data
Oscillated MC w/ syst. error

$$\chi^2 = \sum_{\nu, \bar{\nu}} \sum_i \left[ N^i - \left\{ 1 \pm \frac{1}{2} f_{\nu/\bar{\nu}} \right\} \cdot \left( (1 + f_{\text{sig}}) \cdot n_{\text{sig}}^i + (1 + f_{\nu_\mu}) \cdot n_{\nu_\mu}^i + (1 + f_{\nu_e}) \cdot n_{\nu_e}^i \right) \right]^2 / N^i$$

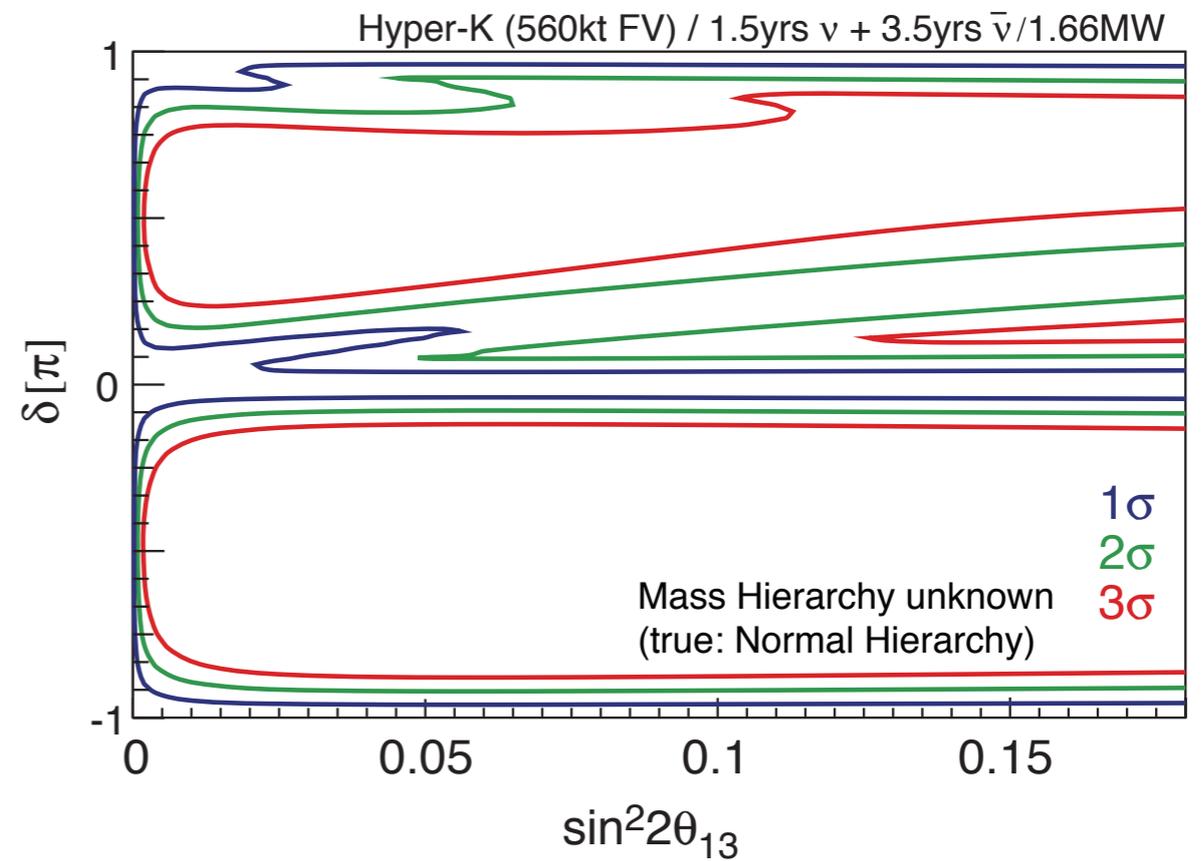
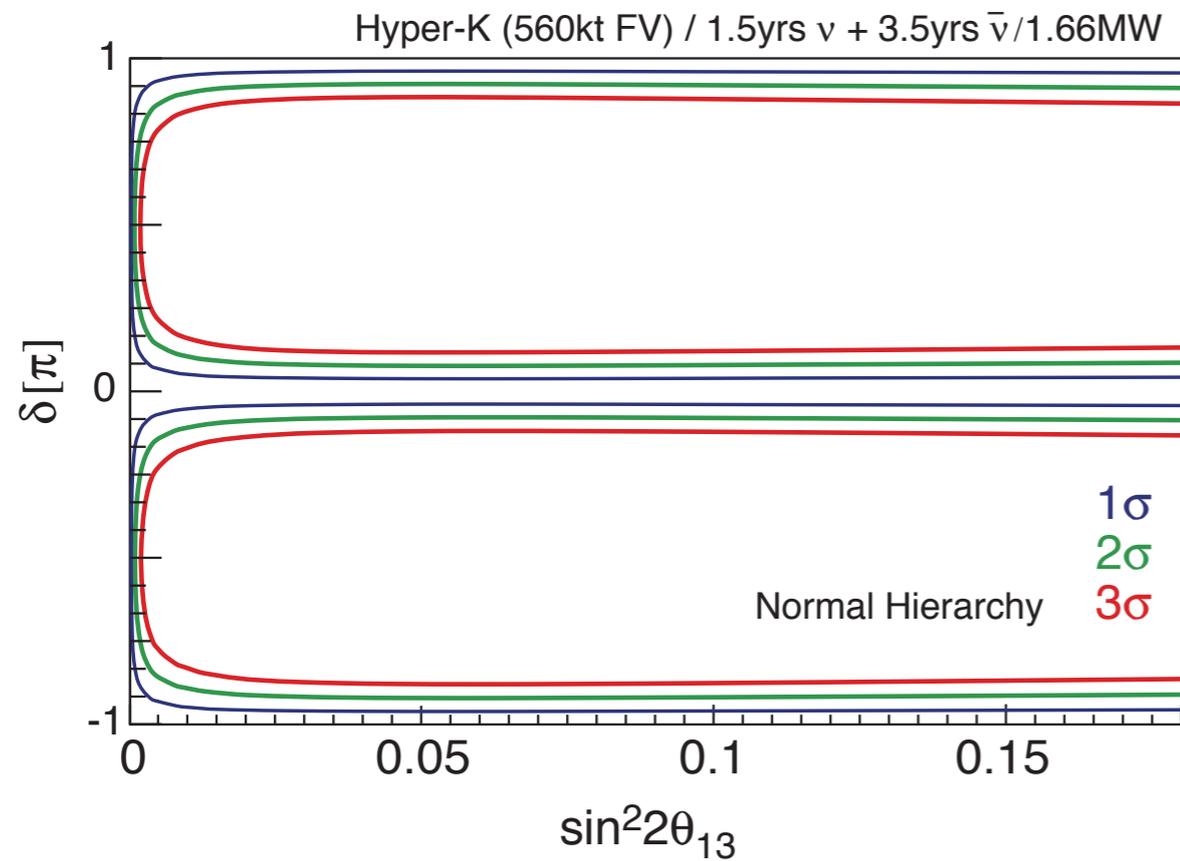
$$+ \frac{f_{\text{sig}}^2}{\sigma_{\text{sig}}^2} + \frac{f_{\nu_\mu}^2}{\sigma_{\nu_\mu}^2} + \frac{f_{\nu_e}^2}{\sigma_{\nu_e}^2} + \frac{f_{\bar{\nu}/\nu}^2}{\sigma_{\bar{\nu}/\nu}^2},$$

assumed syst. errors (all 5%)  
 signal eff.,  $\nu_\mu$ BG,  $\nu_e$ BG,  $\nu/\text{anti-}\nu$  ratio

# $1\sigma$ error of $\delta$

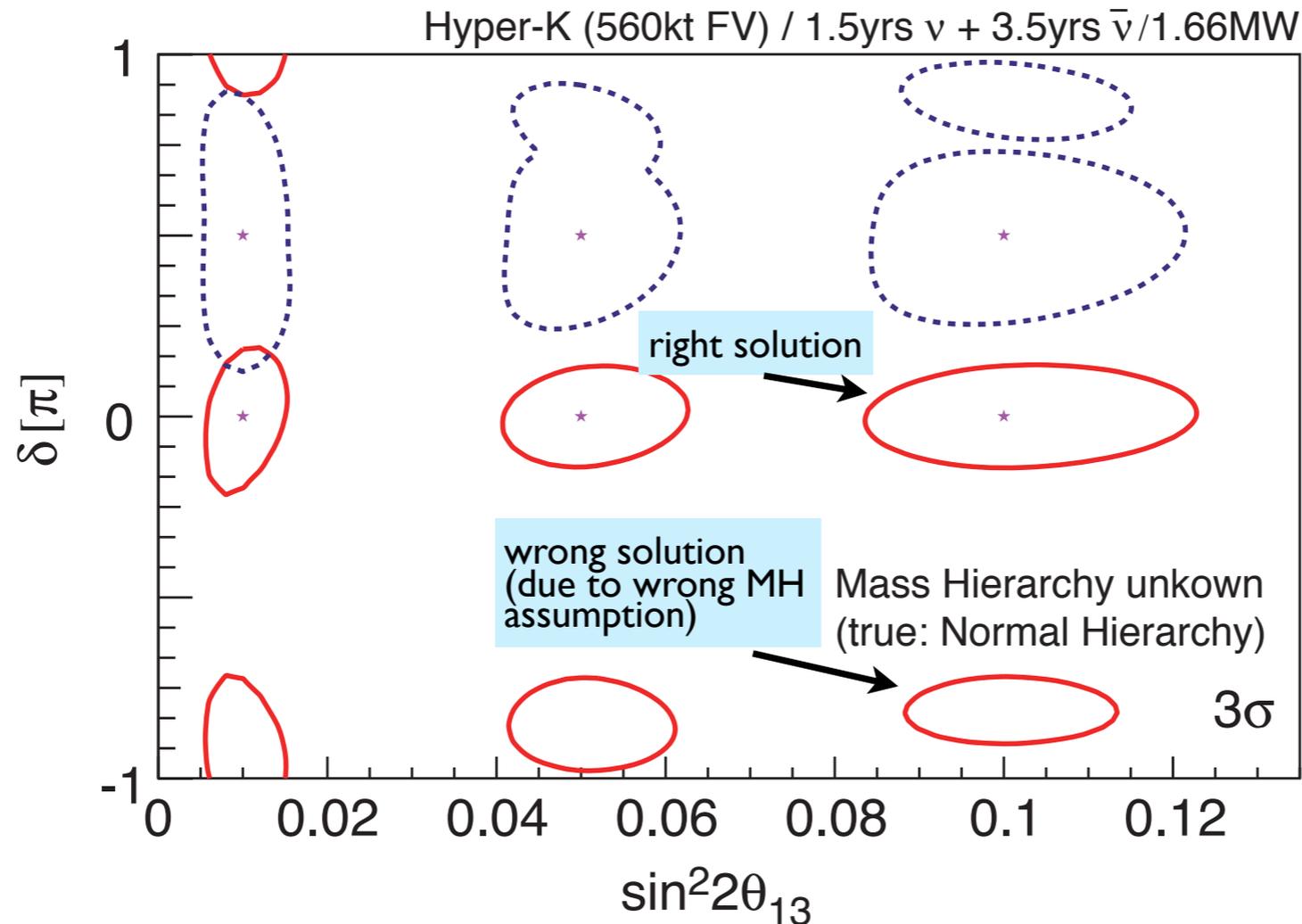


# $\delta$ coverage for CPV



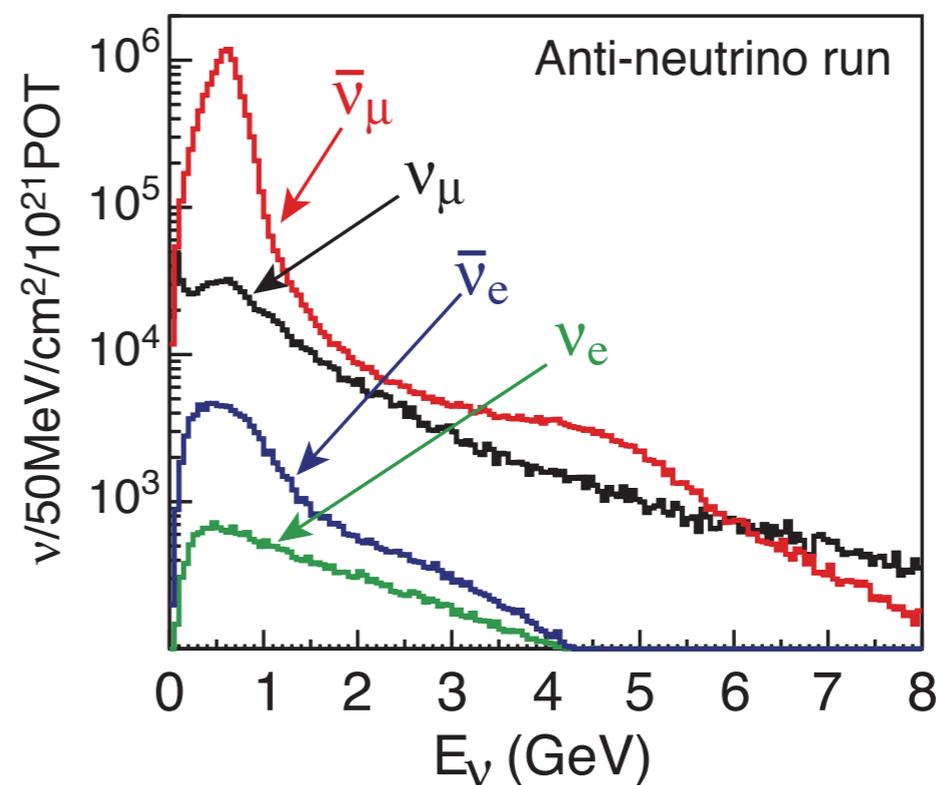
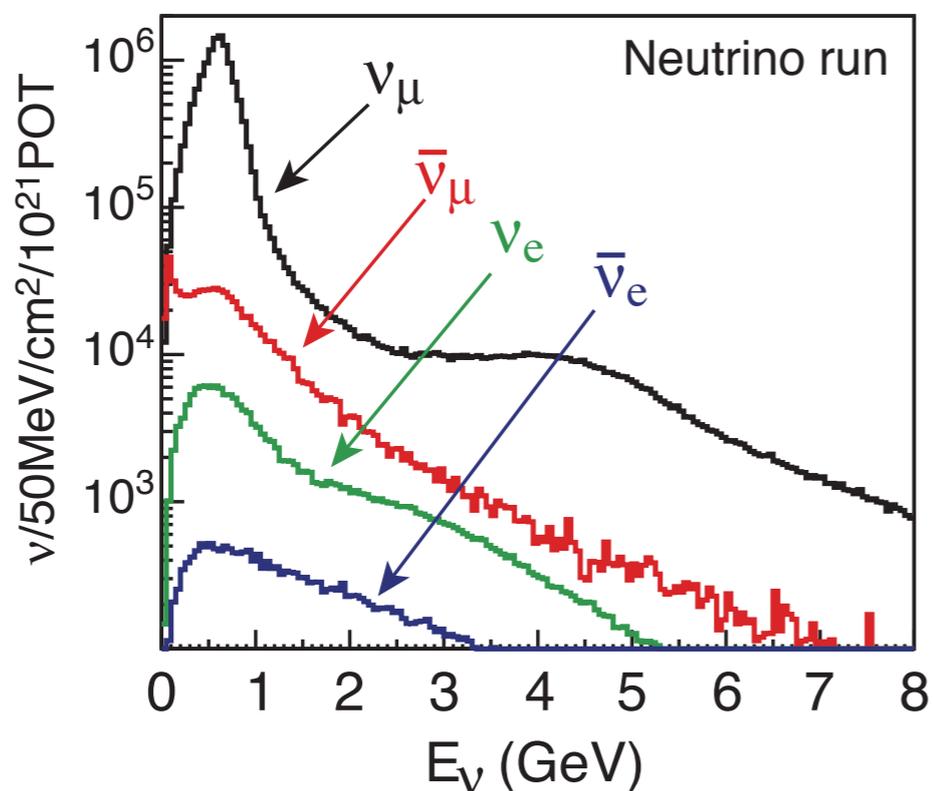
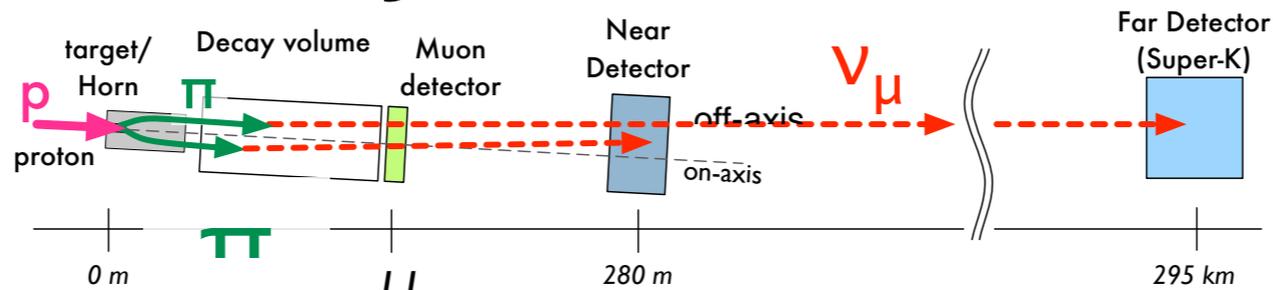
# If mass hierarchy is unknown

Normal mass hierarchy (unknown)



- ▶ multiple solutions, wider allowed region due to wrong MH assumption.
- ▶ Input (mass hierarchy) from other experiments may become important.
  - ▶ from NoVa? or  $\nu$ -less DB? or...
- ▶ One possibility is to determine MH by atm.  $\nu$  study (discuss later)

# JPARC $\nu$



## ▶ J-PARC $\nu$ 's

- ▶ T2K flux based on beamline geometry, horn (320kA), hadron ( $\pi$ , K) production data, proton profile
- ▶ anti- $\nu$  by inverted horn current

- ▶ Quasi-monochromatic  $\nu$  w/ peak energy at oscillation maximum
- ▶ intrinsic beam  $\nu_e$  (BG) is  $<1\%$  at peak energy

# MR Power Improvement Scenario

	Day1 (up to Mar.2011)	Achieved!	Next Step	KEK Roadmap
Power(MW)	0.145		0.45	>1.66
Energy(GeV)	30		30	30
Rep Cycle(sec)	3.04		2.2	1.92~0.5
No. of Bunch	8		8	8
Particle/Bunch	$1.2 \times 10^{13}$		$2.5 \times 10^{13}$	$4.1 \sim 8.3 \times 10^{13}$
Particle/Ring	$9.2 \times 10^{13}$		$2.0 \times 10^{14}$	$3.3 \sim 6.7 \times 10^{14}$
LINAC(MeV)	181		181	400
RCS	h=2		h=2	h=2 or 1

**Combination of High rep. cycle and High beam density**

# Oscillated $\nu_e$ flux

NuclPhysB680,479(2004)

$$\frac{\Phi(\nu_e)}{\Phi_0(\nu_e)} - 1 \approx P_2(r \cdot \cos^2 \theta_{23} - 1) \quad \text{Solar term}$$

$$-r \cdot \sin \tilde{\theta}_{13} \cdot \cos^2 \tilde{\theta}_{13} \cdot \sin 2\theta_{23} (\cos \delta \cdot R_2 - \sin \delta \cdot I_2)$$

$$+2 \sin^2 \tilde{\theta}_{13} (r \cdot \sin^2 \theta_{23} - 1)$$

Interference term ( $\delta$ CP)  
 $\theta_{13}$  resonance term

$r$  :  $\mu/e$  flux ratio ( $\sim 2$  at low energy)

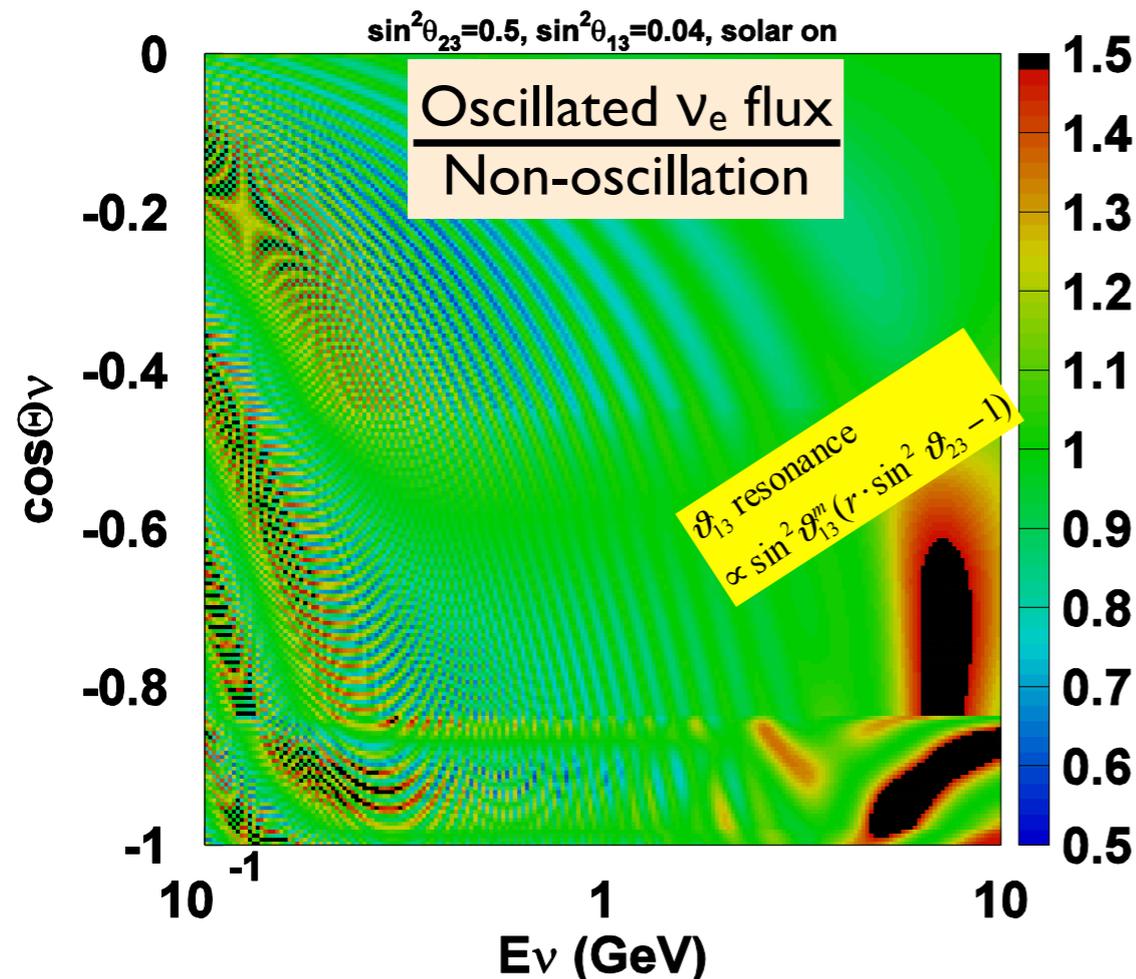
$P_2 = |A_{e\mu}|^2$ :  $2\nu$  transition probability  $\nu_e \rightarrow \nu_{\mu\tau}$  in matter

$R_2 = \text{Re}(A_{ee}^* A_{e\mu})$

$I_2 = \text{Im}(A_{ee}^* A_{e\mu})$

$A_{ee}$ : survival amplitude of the  $2\nu$  system

$A_{e\mu}$ : transition amplitude of the  $2\nu$  system



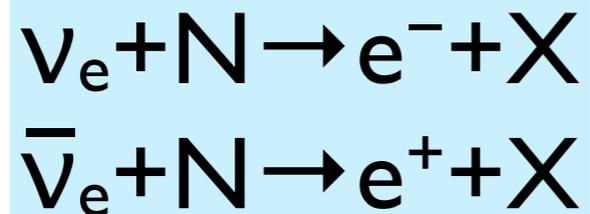
## Relevant $\nu$ oscillation parameters

- ▶ Solar term
  - ▶  $\sin^2 \theta_{23}$
- ▶ Interference term
  - ▶  $\delta$ , matter effect ( $\theta_{13}$ , Mass Hierarchy)
- ▶  $\theta_{13}$  resonance term
  - ▶  $\sin^2 \theta_{23}$ , matter effect ( $\theta_{13}$ , Mass Hierarchy)

Effective  $\theta_{13}$  becomes large due to Earth's matter potential  $\Rightarrow$  more  $\nu_e$  appearance

- happens in  $\nu$  in the case of normal mass hierarchy
- in anti- $\nu$  in inverted mass hierarchy

# $\nu_e$ -like and anti- $\nu_e$ -like sample



Upward  $\nu_e$  appearance

- ▶  $\nu_e$  CC produce more  $\pi^+$ 
  - ▶ more muon decays
- ▶ More energy transfer to hadronic system
  - ▶ lower charged lepton energy
  - ▶ more pions (muon decays)



enrichment by variables

- ▶ # of rings
- ▶ # of muon decay electrons
- ▶ Lepton energy fraction
- ▶ transverse momentum

	$\nu_e$ CC	anti- $\nu_e$ CC	others	Total
$\nu_e$ -like	57%	11%	32.0%	100%
anti- $\nu_e$ -like	55%	34%	11%	100%

$\sin^2 2\theta_{13} = 0.1$

